

Track direction identification via track-fitting quality for cosmic-ray background suppression in the COMET experiment



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Contents

See also:

M. Moritsu et al., [arXiv:2508.15344](https://arxiv.org/abs/2508.15344)

▶ Introduction

- The COMET experiment & the tracking system

▶ Motivation

- Cosmic-ray background in the COMET experiment

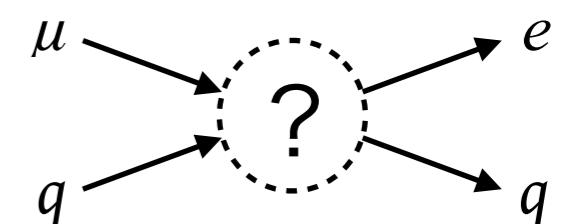
▶ Track-direction ID method

- to eliminate reverse-traveling opposite-sign particles

μ -e conversion & COMET

► Muon-to-electron conversion:

- Neutrinoless coherent transition in a nuclear field
- Violates the Lepton Flavor conservation
- Search for NEW physics beyond the Standard Model



► The COMET experiment:

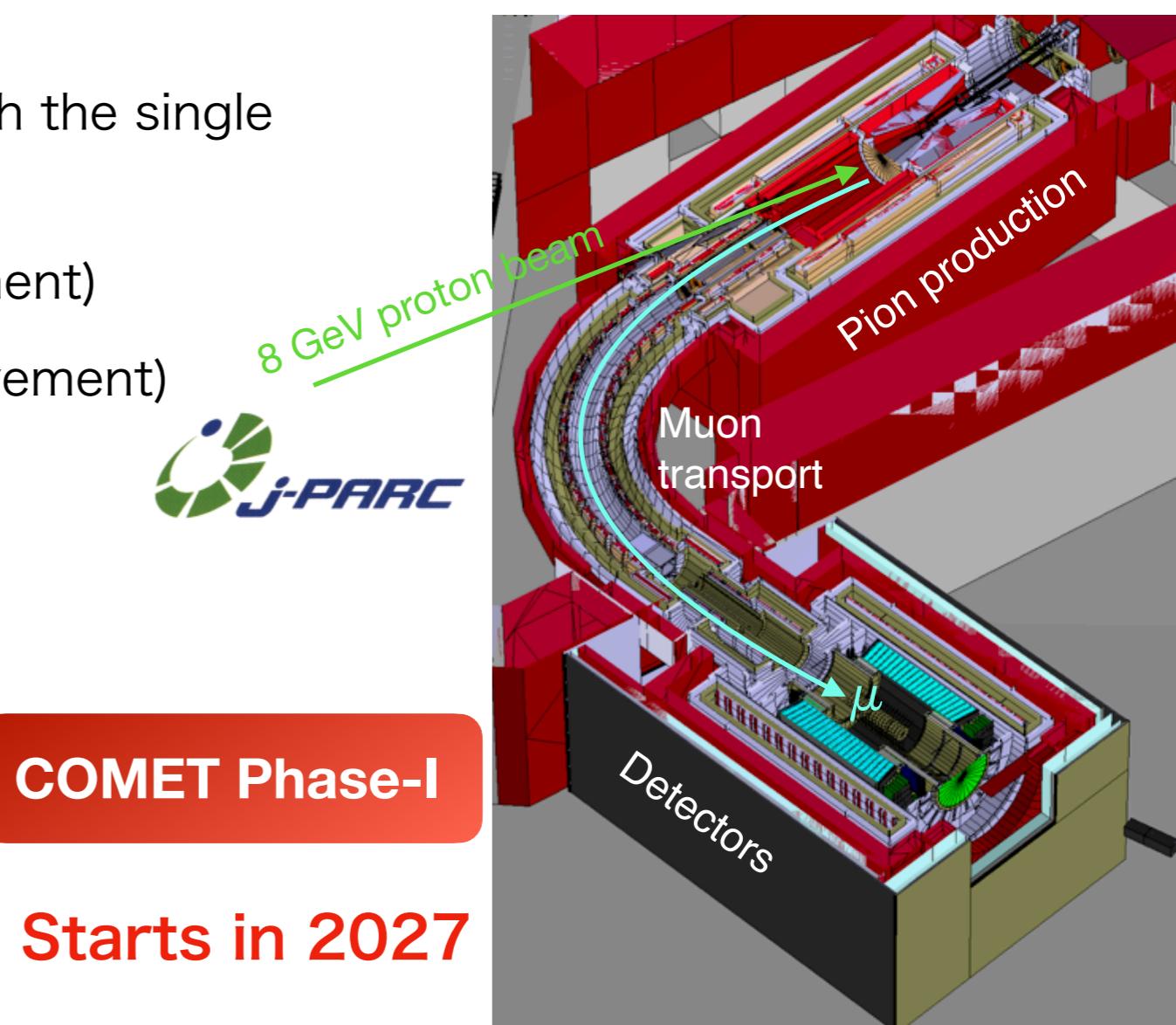
- Explores μ -e conversion at **J-PARC** with the single event sensitivity of
 - Phase-I: **3×10^{-15}** ($\times 100$ improvement)
 - Phase-II: **3×10^{-17}** ($\times 10,000$ improvement)



COMET Phase-I

Technical Design Report
PTEP 2020, 033C01

Starts in 2027

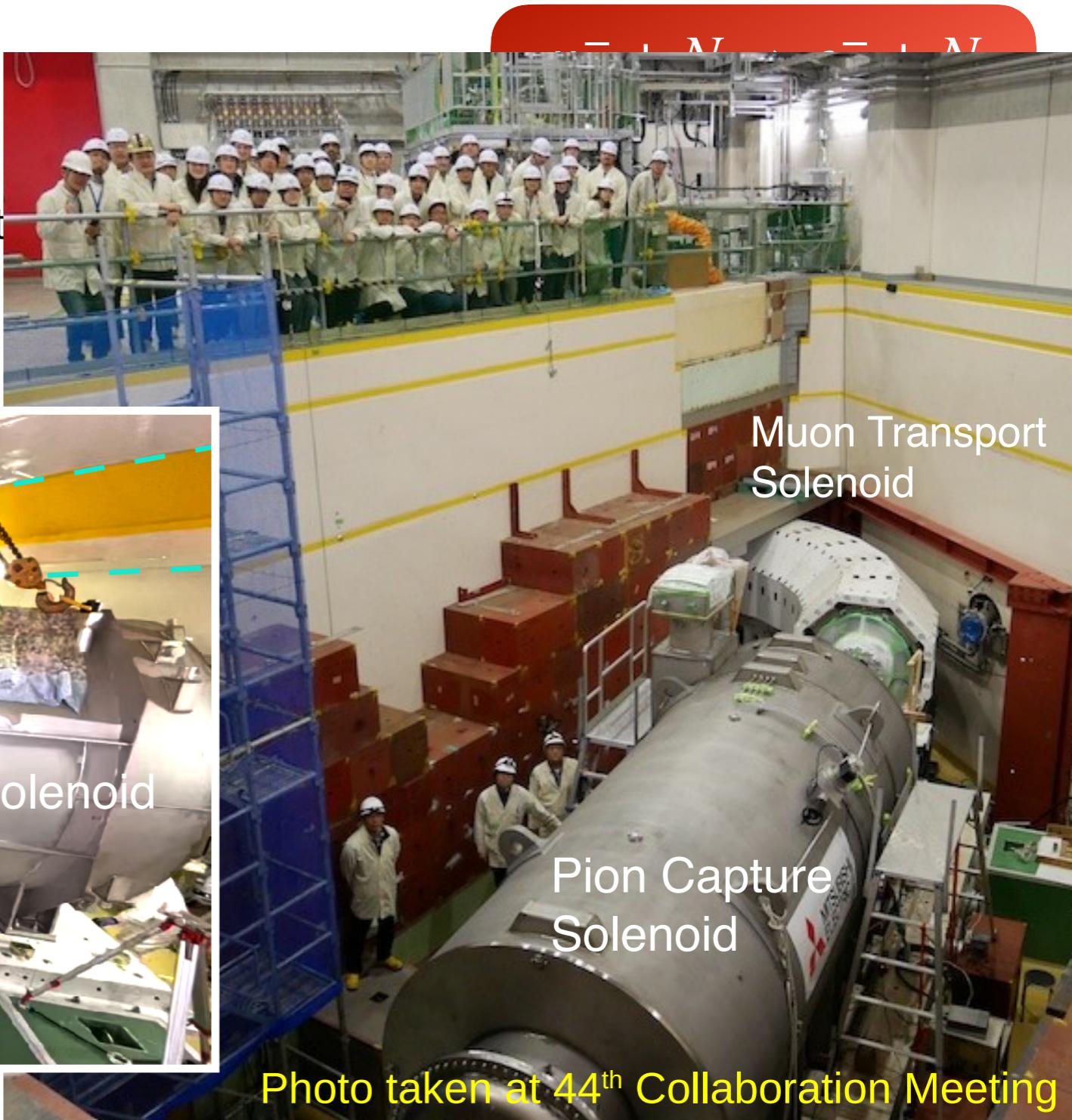


μ -e conversion & COMET

► Muon-to-electron conversion:

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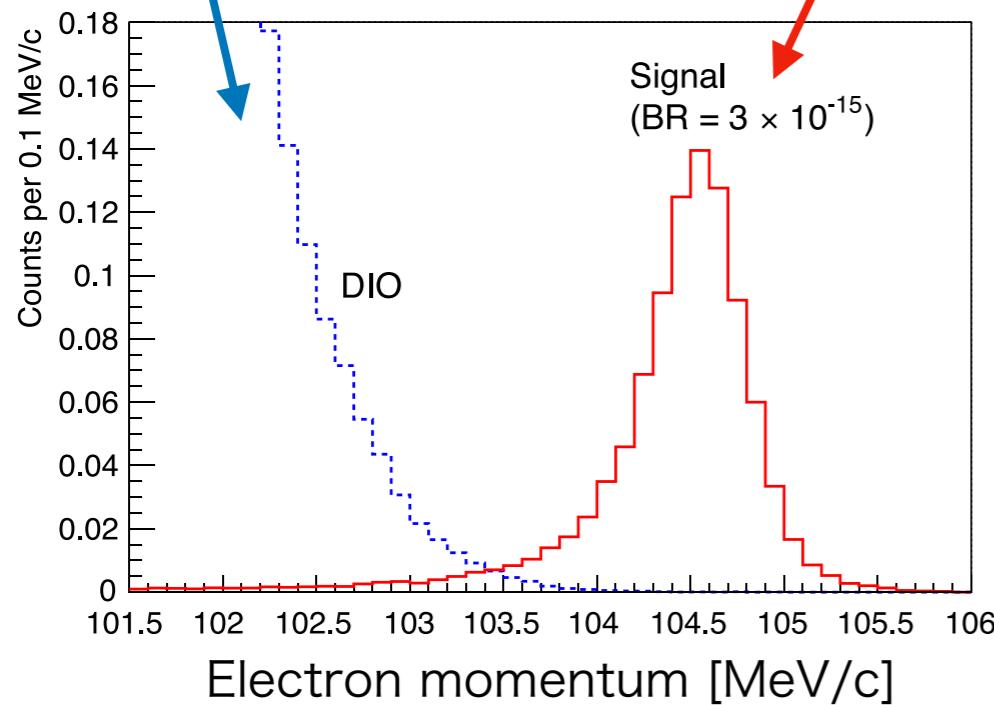
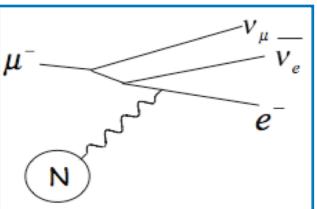
► The COMET experiment:



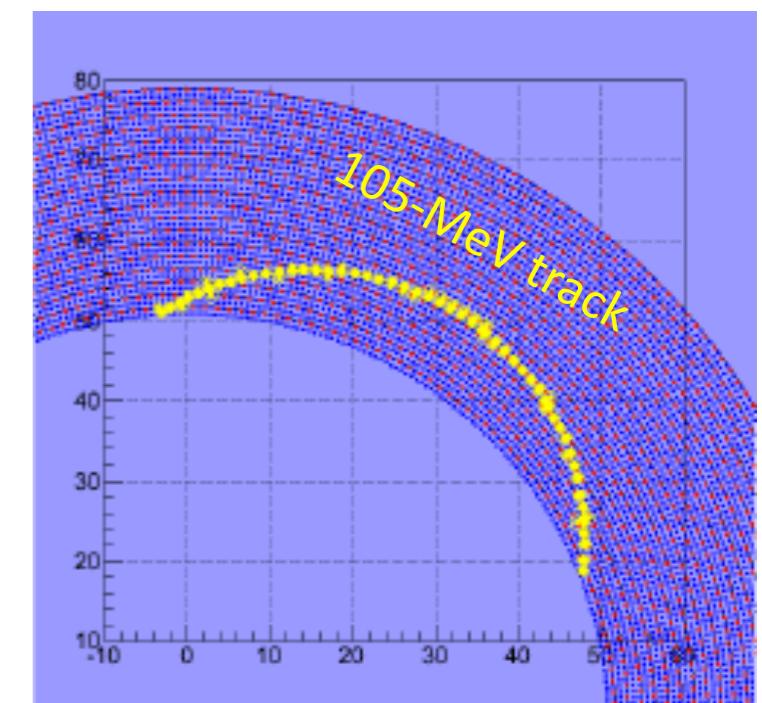
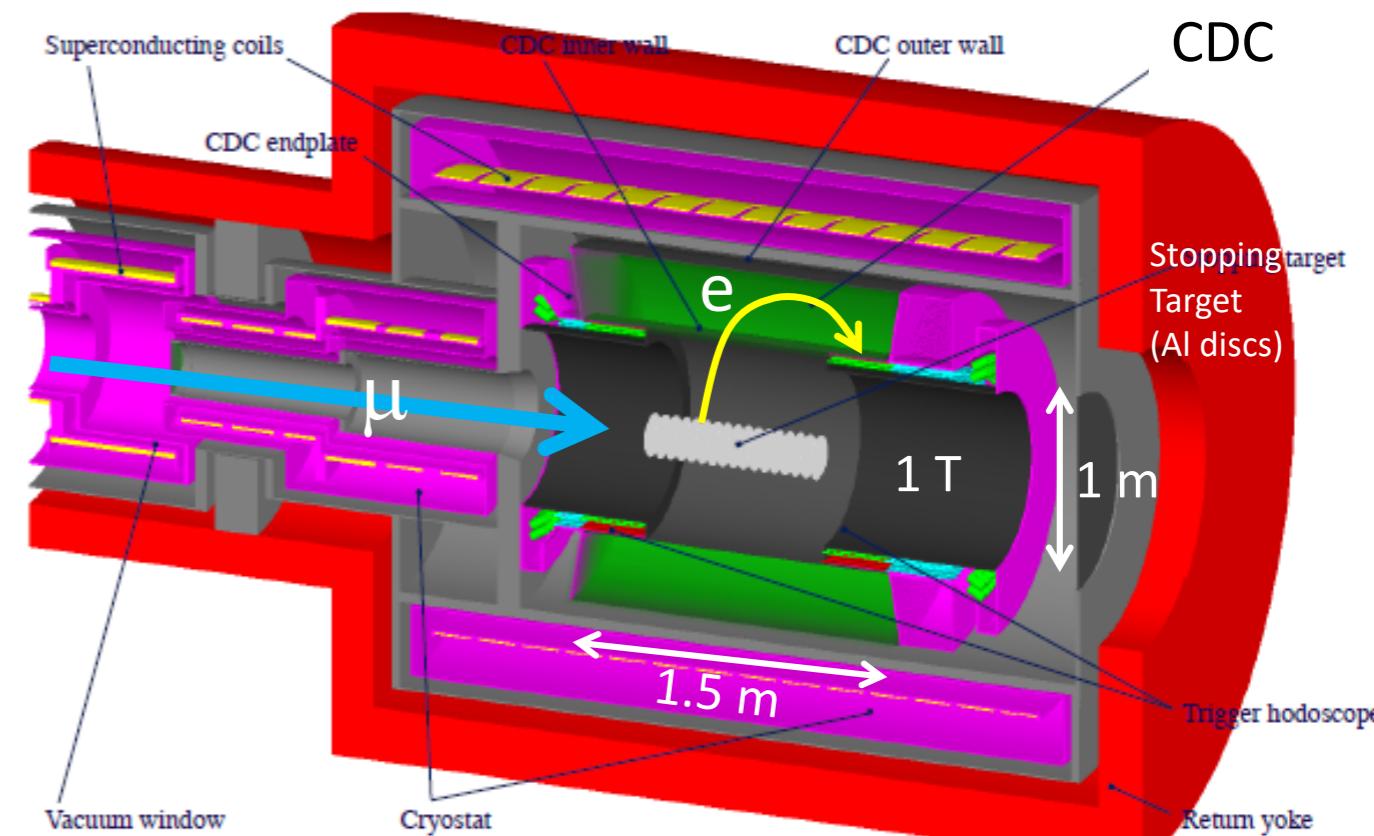
COMET CDC

A. Sato et al., NIM A 1069, 169926 (2024)

Decay-In-Orbit



μ -e conversion



Cylindrical Drift Chamber (CDC)

- ▶ 20 layers of alternating all-sereo wires
- ▶ Low mass gas mixture = He:iC₄H₁₀ (90:10)

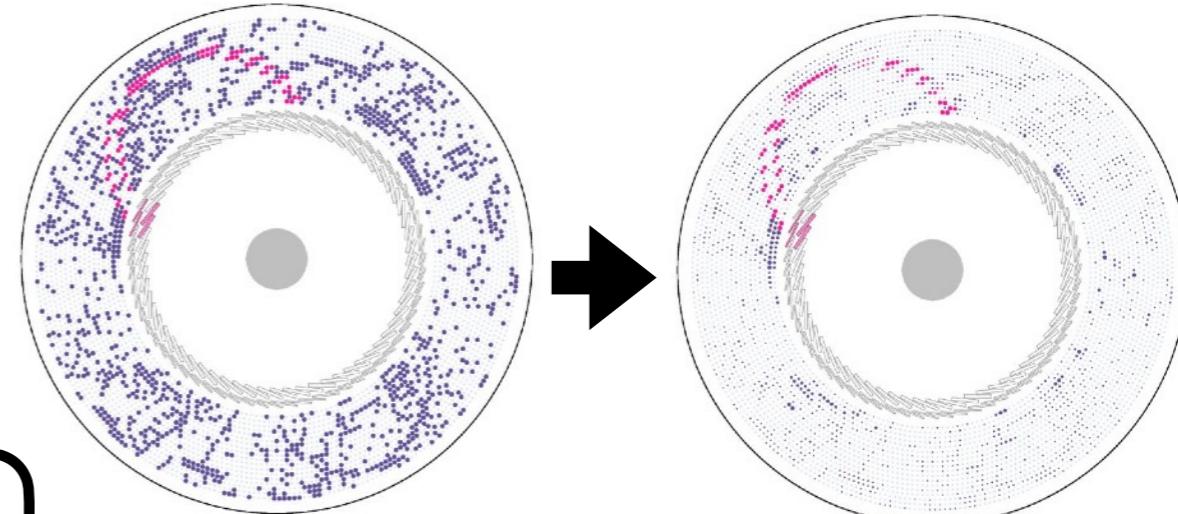
Cylindrical Trigger Hodoscopes (CTH)

- ▶ 256 segmented plastic scintillators

Track Reconstruction

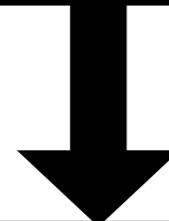
Hit Classification

By using GBDT, 98% of noises are rejected while keeping 99% of signals.



Track Finding

Hough Transformation, Deep Learning, RANSAC, etc.



[F. Kaneko et al., PTEP 2025, 053C01](#)

see also a poster by T. Xing

Track Turn Identification

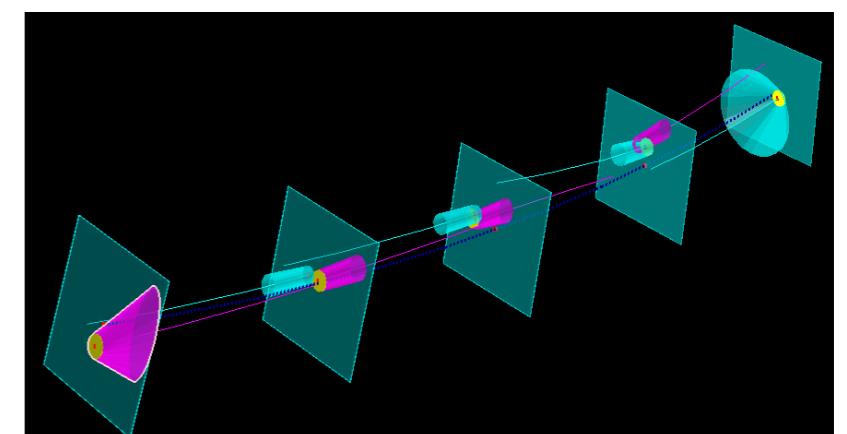
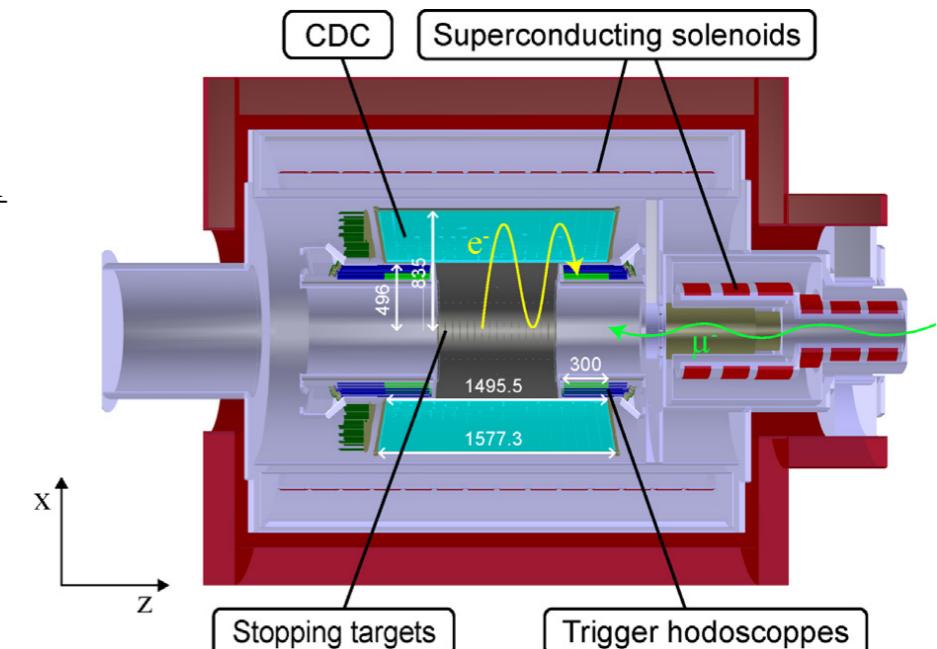
32% of signal events are multiple turn.

We will focus on only single turn in this report.



Track Fitting

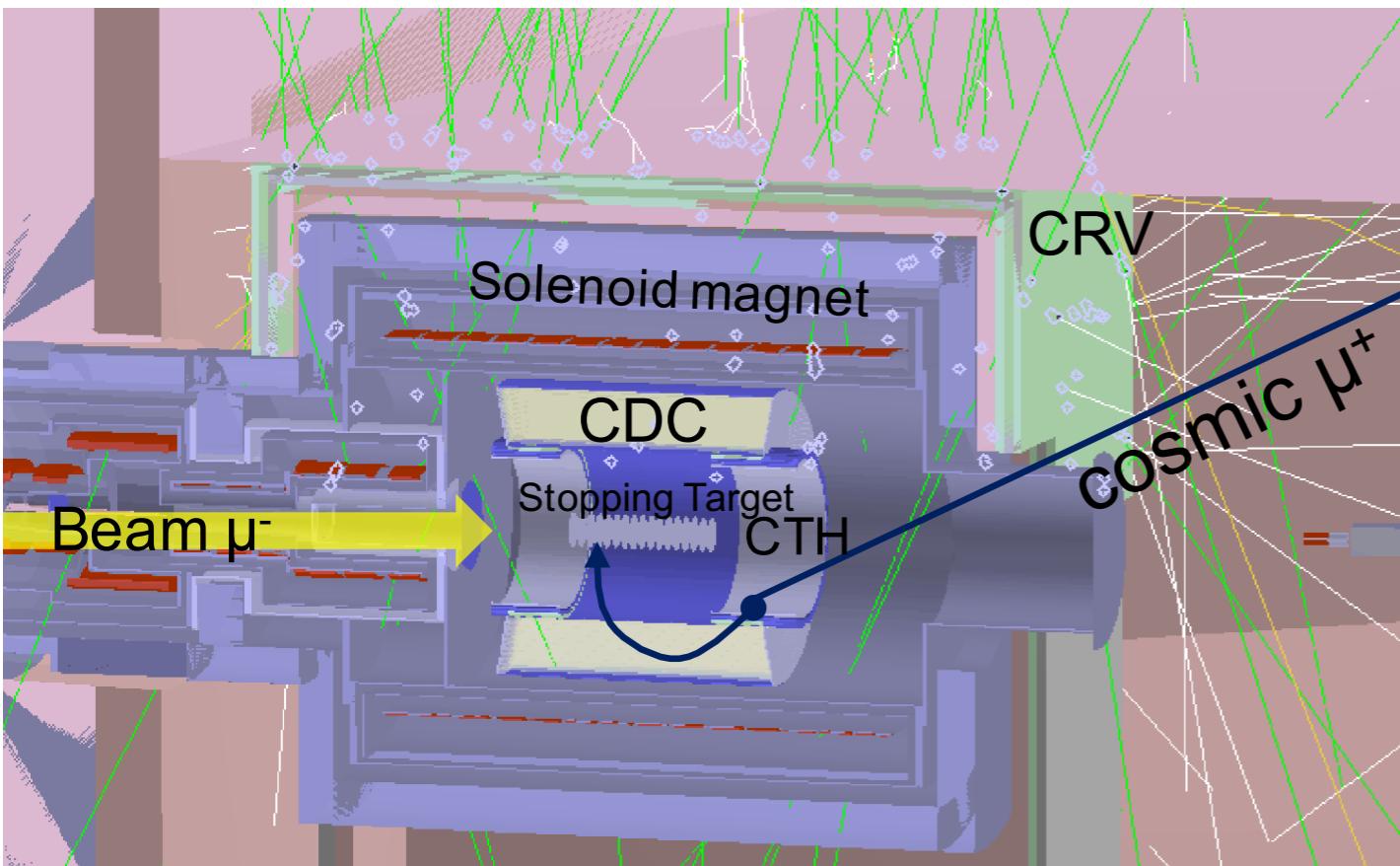
We use Deterministic Annealing Filter in GENFIT2



Motivation

Cosmic-ray background in the COMET experiment

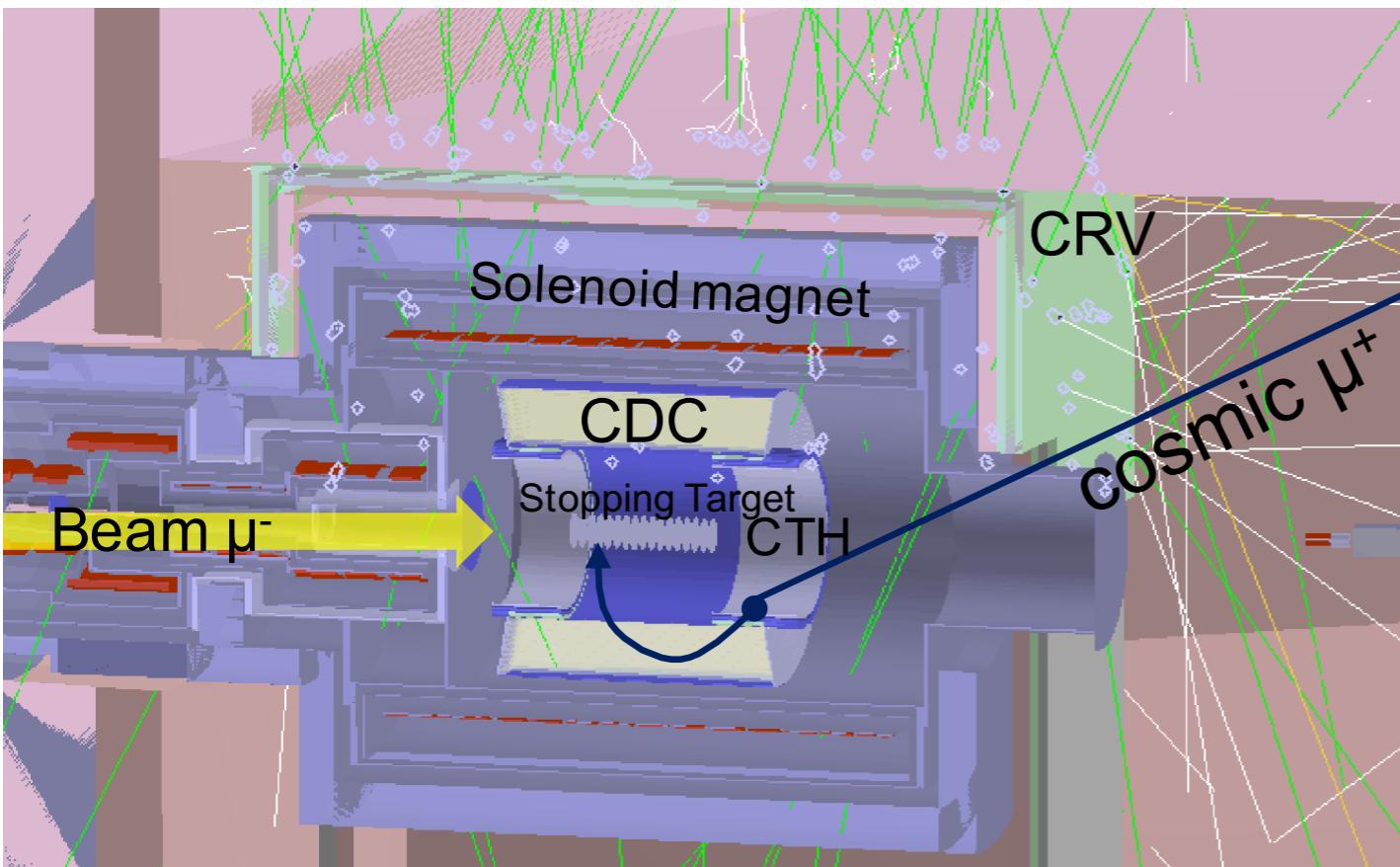
Motivation

 — Cosmic-ray BG

- ▶ **Cosmic rays** may create 105-MeV electrons that come into a detector and make trigger.
- ▶ Cosmic-ray induced BG is basically eliminated by **veto counter** arrays which cover Detector Solenoid (DS) Magnet, ✓ with 99.99% veto efficiency.
- ▶ However, there are uncovered holes through which cosmic rays **may sneak into** DS.

Motivation

— Sneaking cosmic-ray BG



Notice Why μ^+ (not μ^-)

The trajectory of positive charged particles with reverse track direction looks the same as the signal electron track.

Sneaking cosmic μ^+ can be a significant BG.

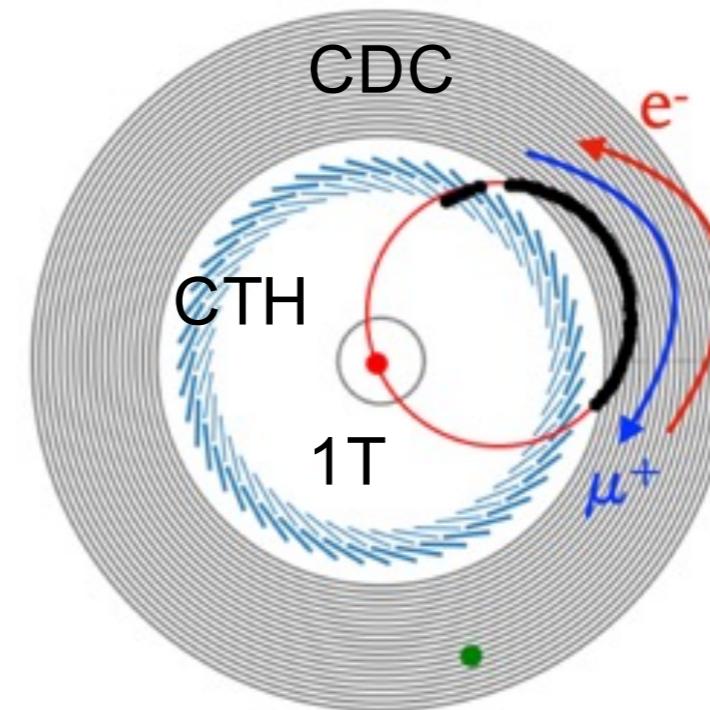
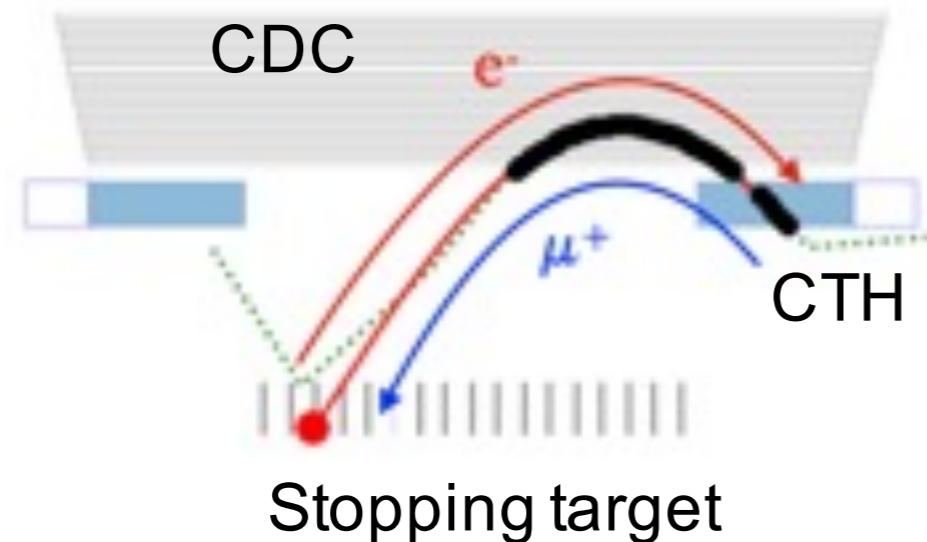
- does not hit CRV (sneaking from downstream hole of DS),
- scattered by CTH support frame,
- hits CTH and enters CDC from the reverse direction compared to the signal electron track, and
- finally hits stopping target.

In the latest simulation, without any measure, 2.4 ± 0.9 BG events remains.

→ Next page

Feature of sneaking μ^+ BG

	μ -e Conv. Signal	Sneaking cosmic BG
Particle	e^-	μ^+
Speed β	1	0.7
Track direction	Target \rightarrow CDC \rightarrow CTH (Normal)	CTH \rightarrow CDC \rightarrow Target (Reverse)

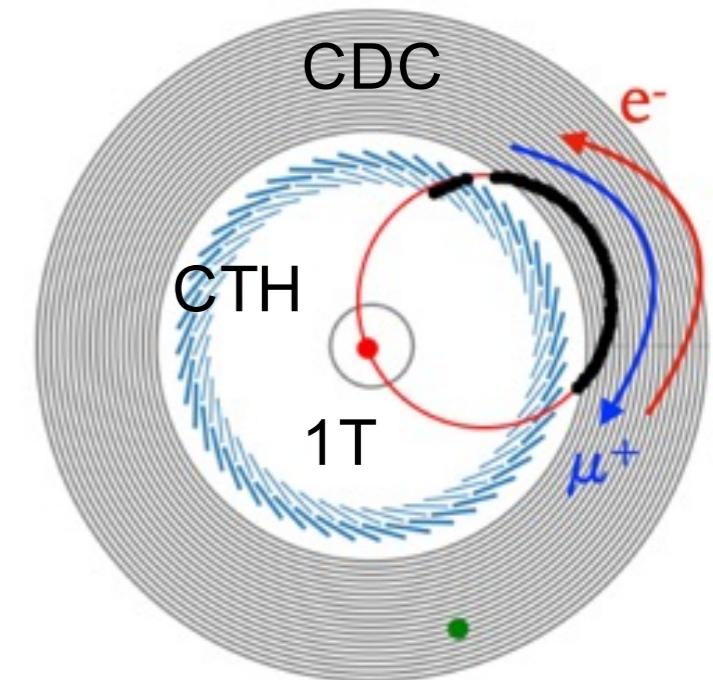


Can we know track direction ?

We normally assume the track direction is from Target to CTH.

TOF miscorrection

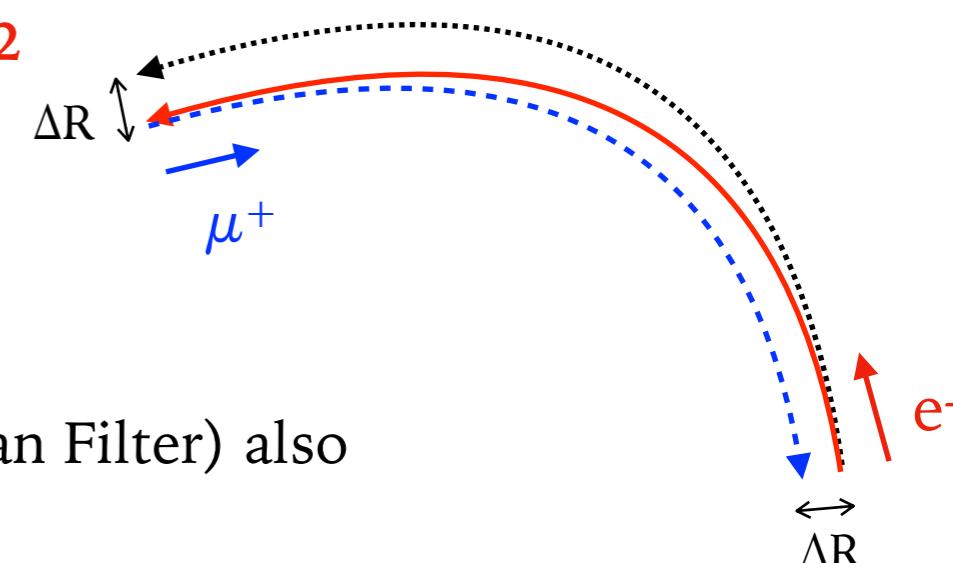
- In order to know correct drift time, **time-0** is corrected according to TOF between CTH and each CDC hit.
- If the track is reverse, this makes miscorrection of **at longest 9.7 ns** for reverse μ^+ .
 - $9.7 \text{ ns} \times 25 \mu\text{m/ns}^* = 240 \mu\text{m}$
 - (comparable to spatial resolution of $150 \mu\text{m}$)



* Drift velocity for He:iC₄H₁₀ (90:10)
 ~ typically $25 \mu\text{m/ns}$

Idea

This miscorrection will make the **difference in χ^2** between normal & reverse direction hypotheses.



Note:

Material effect (energy loss correction in Kalman Filter) also makes χ^2 worse for reverse tracks.

Track-direction ID method

Simulation & Analysis

MC event generation

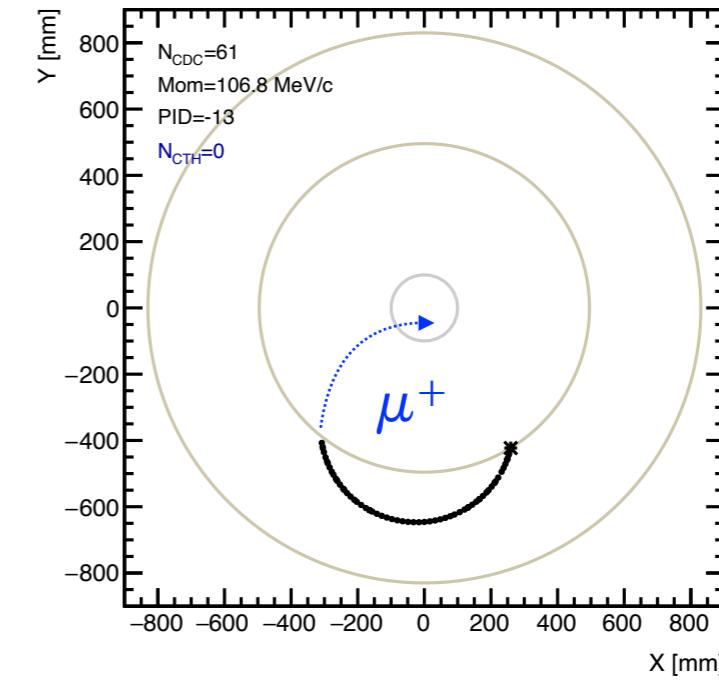
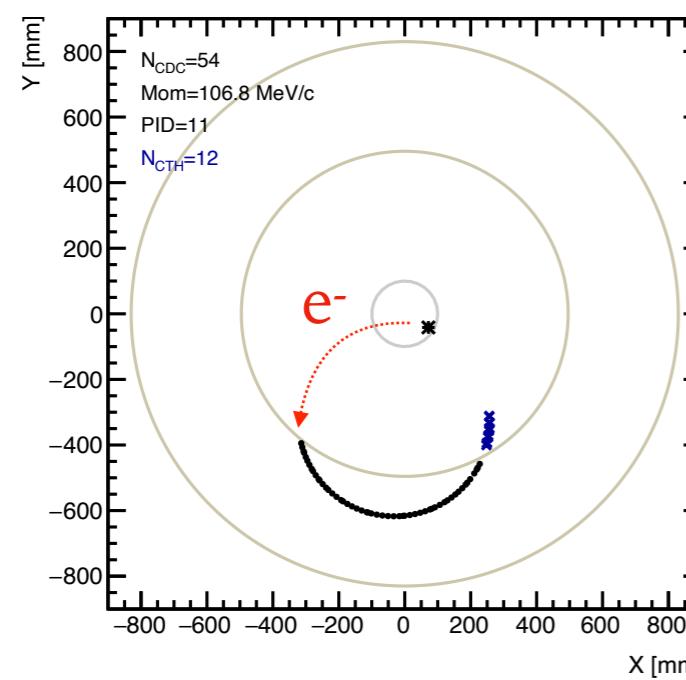
Signal e^- MC sample

- electron e^-
- 100—110 MeV/c
- From stopping target disks
- CDC max layer ≥ 5
- Require CTH hits
- Single turn only
- Spatial resolution: 150 μm

Reversed

Reversed μ^+ MC sample:

- muon μ^+
- 100—110 MeV/c
- From CTH to target disks
- CDC max layer ≥ 5
- Single turn only
- Spatial resolution: 150 μm



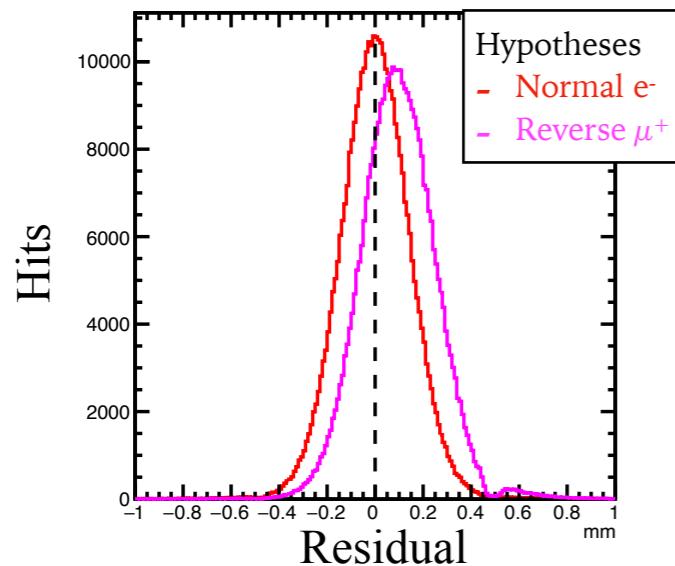
Track Fitting with 2 Hypotheses

Compare 2 hypotheses

Normal-direction e^- vs Reverse-direction μ^+

Signal e^- MC sample

This induces **mis-correction of TOF**



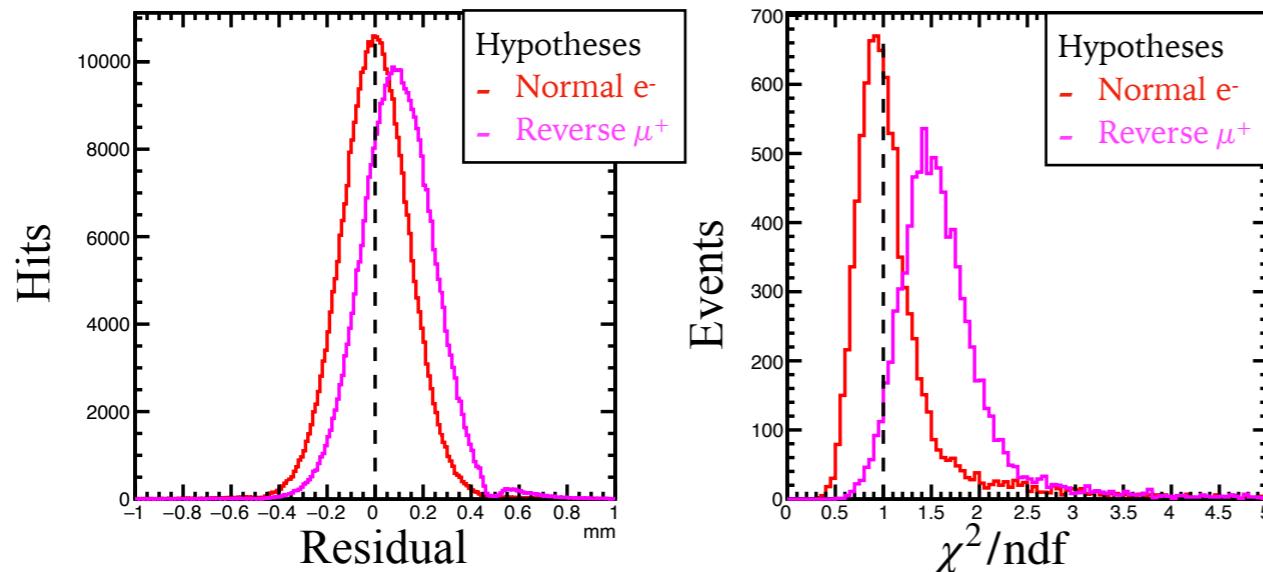
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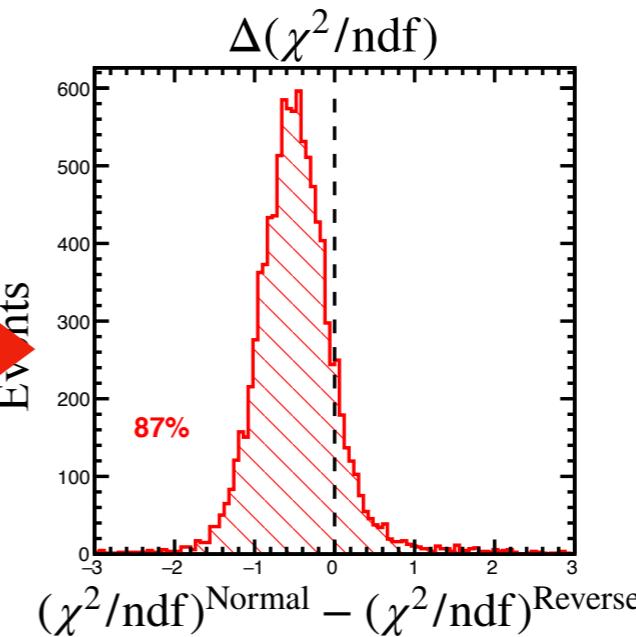
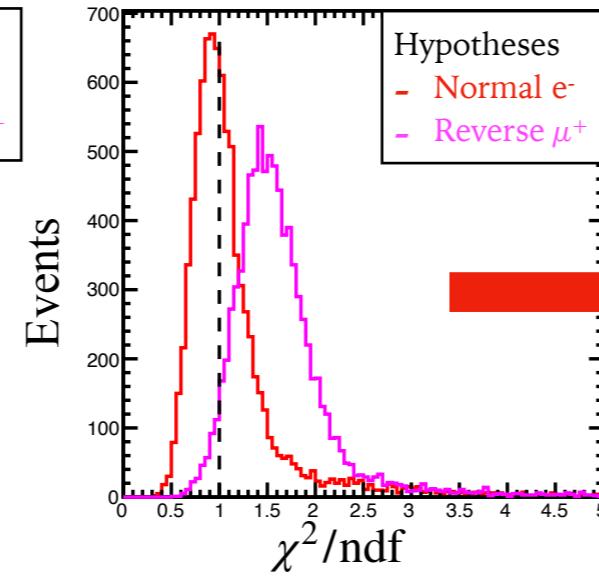
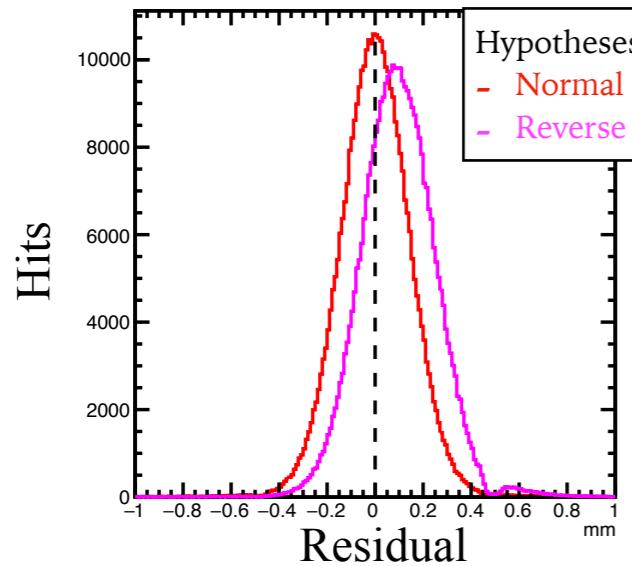
Difference does exist !

Track Fitting with 2 Hypotheses

Compare 2 hypotheses

Normal-direction e^- vs Reverse-direction μ^+

Signal e^- MC sample



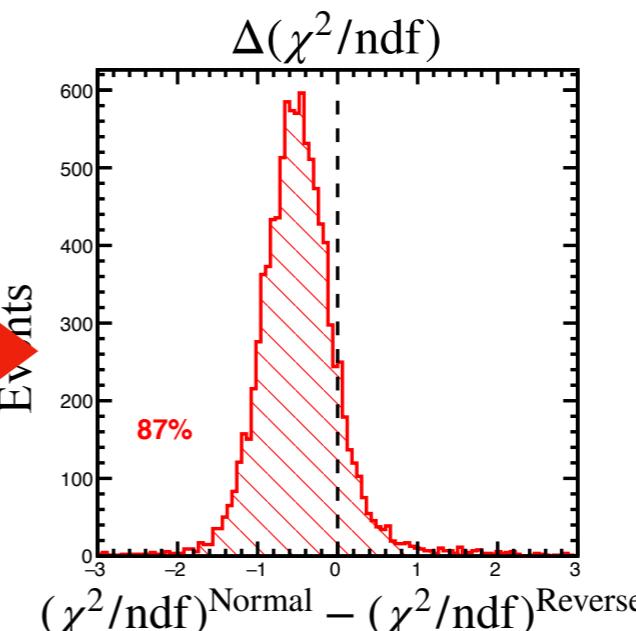
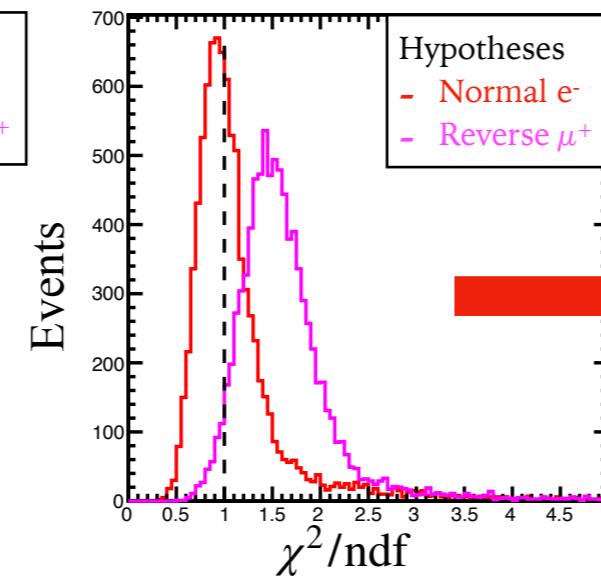
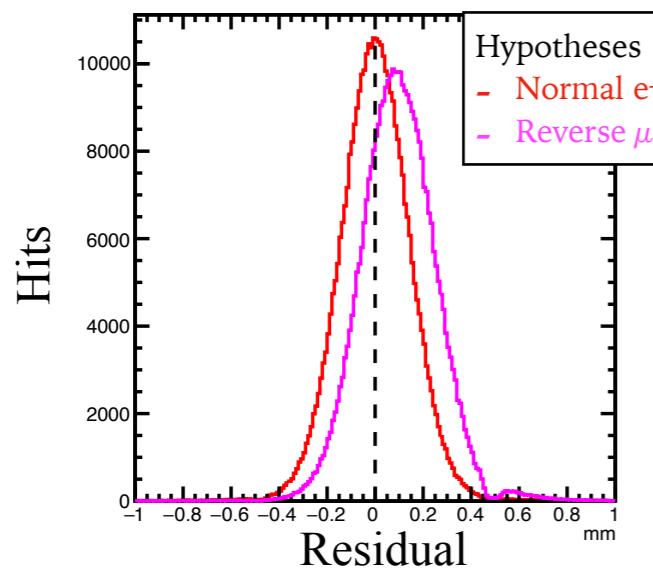
This induces **miscorrection of TOF**

Track Fitting with 2 Hypotheses

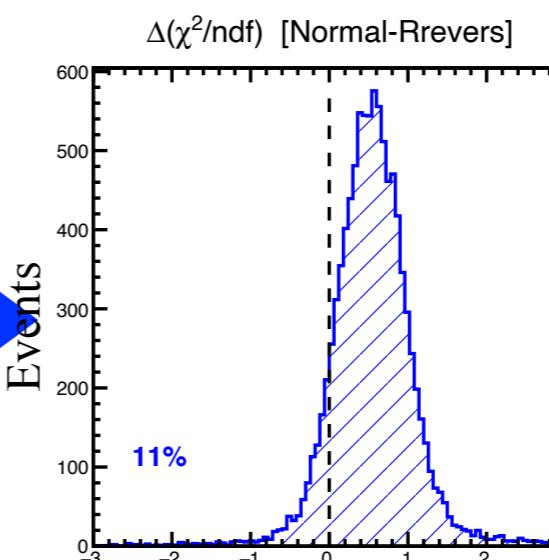
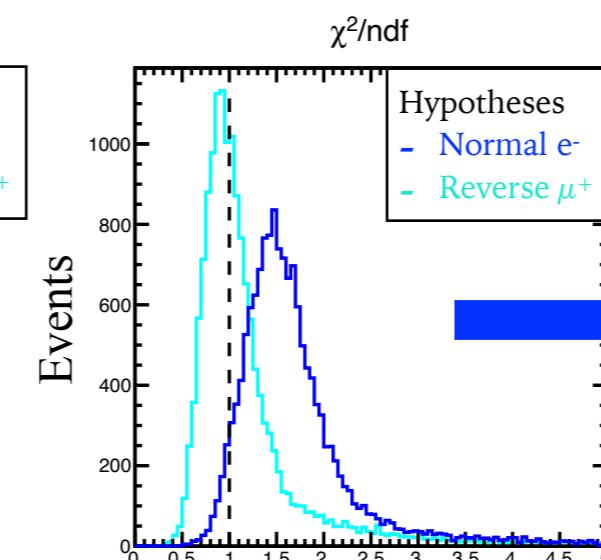
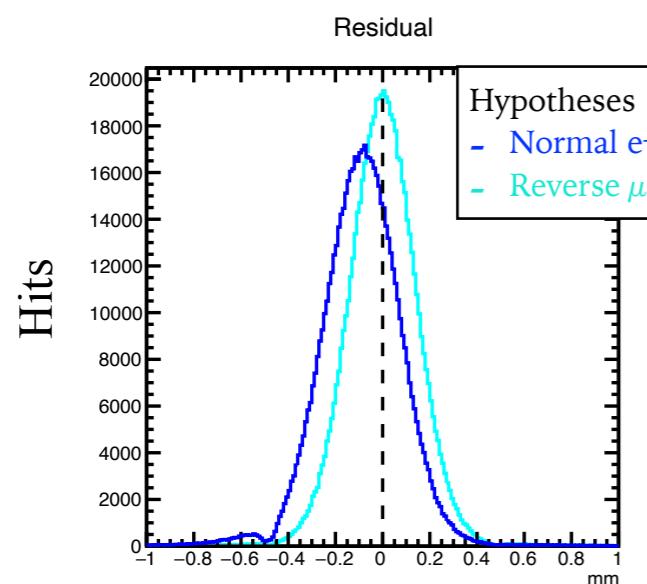
Compare 2 hypotheses

Normal-direction e^- vs Reverse-direction μ^+

Signal e^- MC sample



Reversed μ^+ MC sample

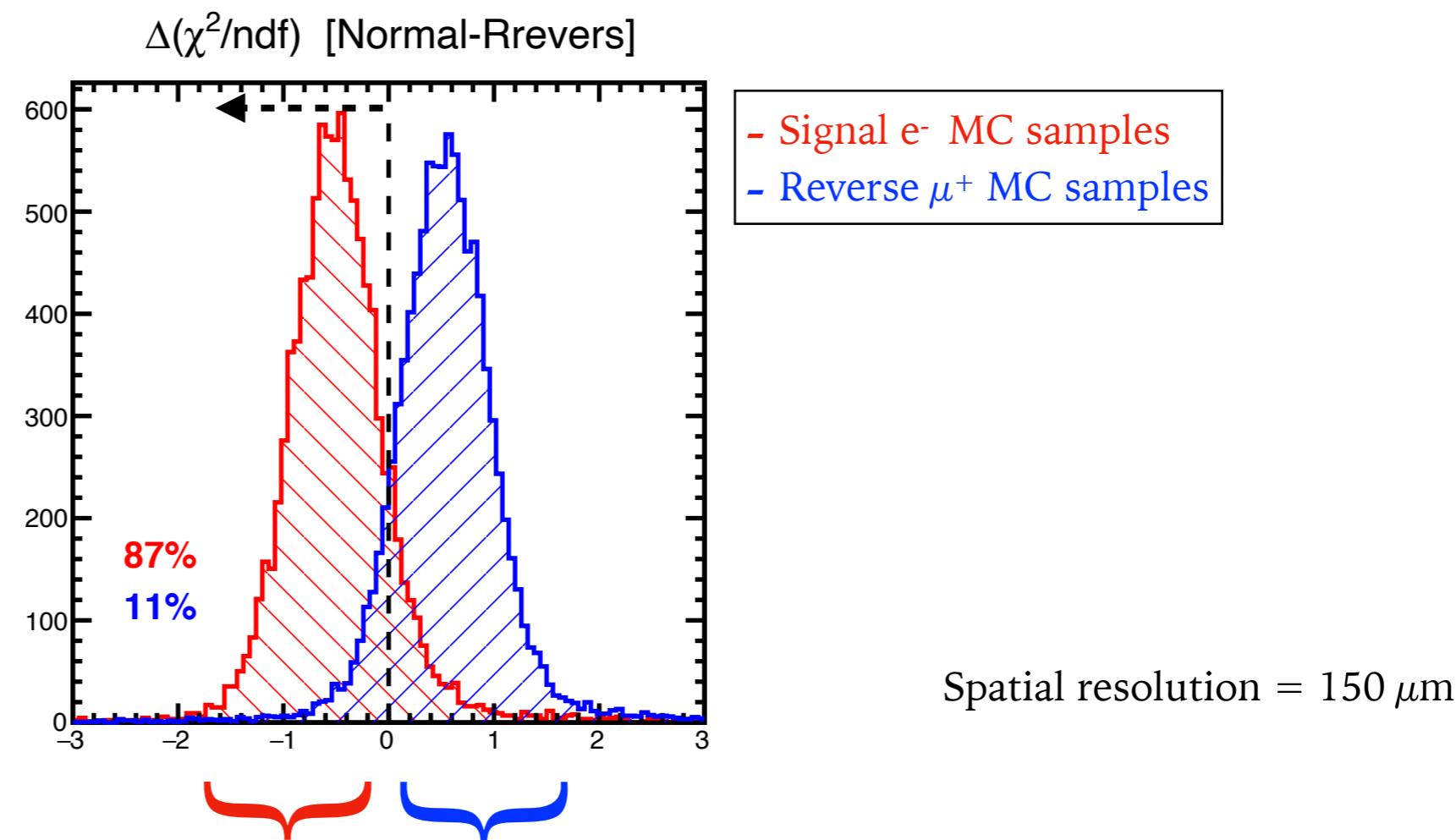


This induces **misCorrection of TOF**

Signal retention & BG rejection



If we set threshold as
 $\Delta(\chi^2/\text{ndf}) < 0$,

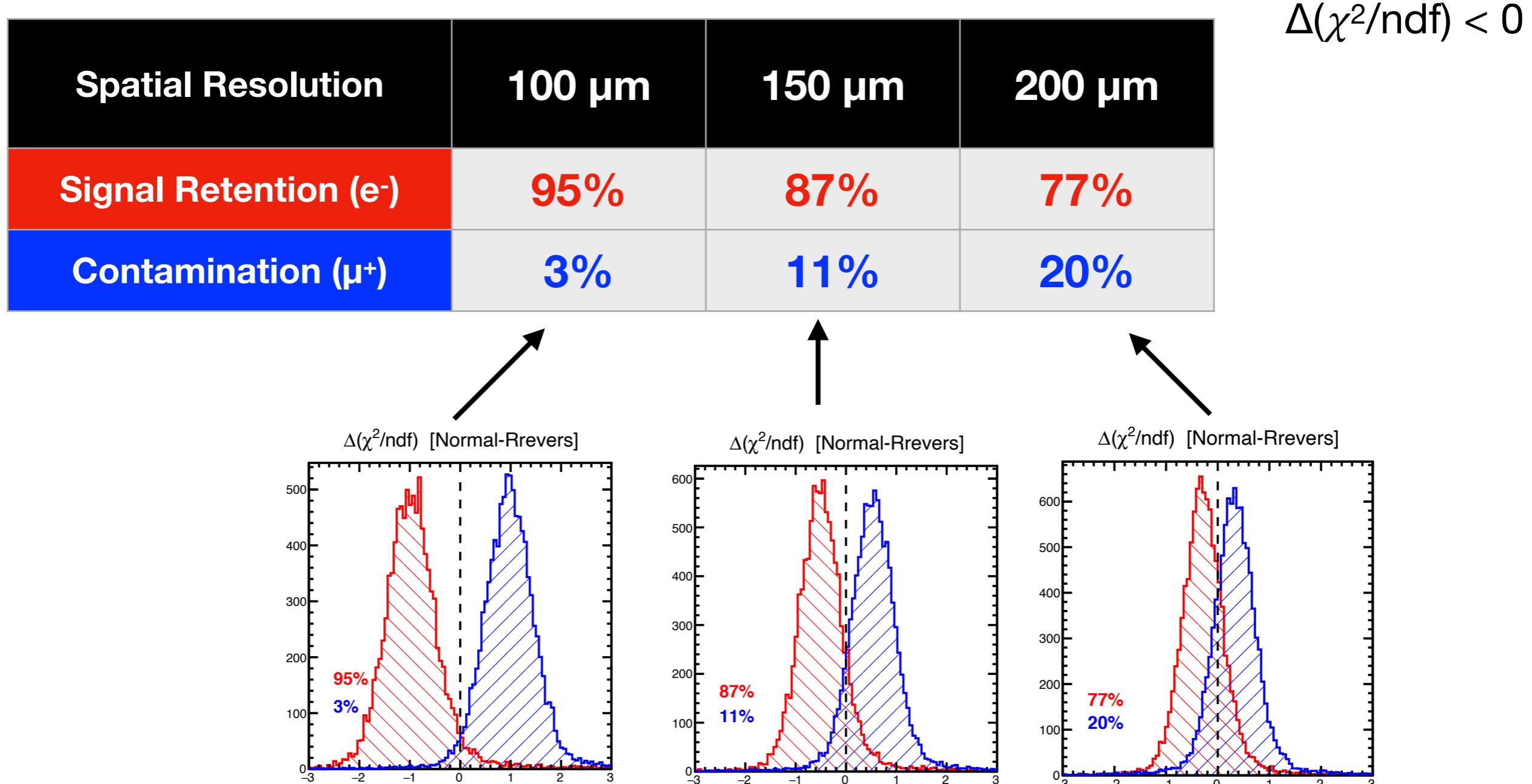


87% of signal e^-
are retained

89% of reverse μ^+
are eliminated

BG can be suppressed to **11%**

Dependence of spatial resolution



✓ Rejection power is sensitive to spatial resolution.

We expect $\sim 150 \mu\text{m}$ resolution based on a prototype test

Reduction of BG events

Spatial Resolution	100 μm	150 μm	200 μm	Without Direction ID method
Signal Retention (e ⁻)	95%	87%	77%	100%
Contamination (μ^+)	3%	11%	20%	100%
BG events	0.07	0.26	0.48	2.4

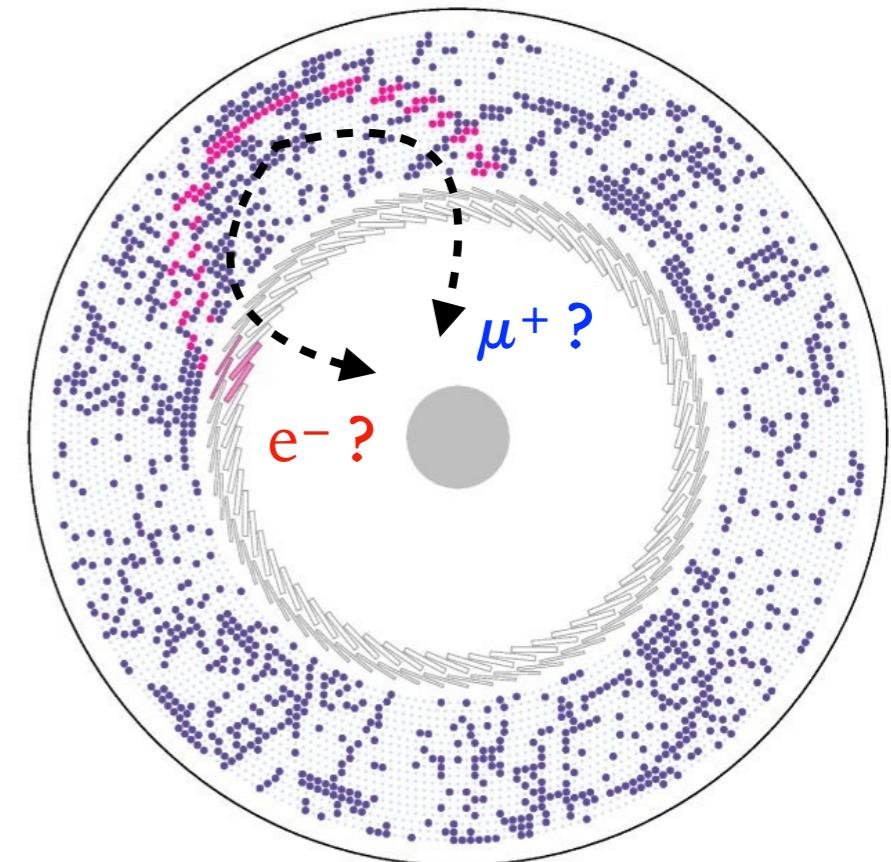
Latest
Prediction



Assuming 150 μm spatial resolution,
Sneaking cosmic μ^+ BG can be reduced to **0.26** events

Next step — Outlook

- ▶ There is room to improve reduction power by using machine learning with multi-variate analysis, instead of simple comparison of χ^2/ndf .
- ▶ We have to check the performance with many noise hits.



Summary

See also:

M. Moritsu et al., [arXiv:2508.15344](https://arxiv.org/abs/2508.15344)

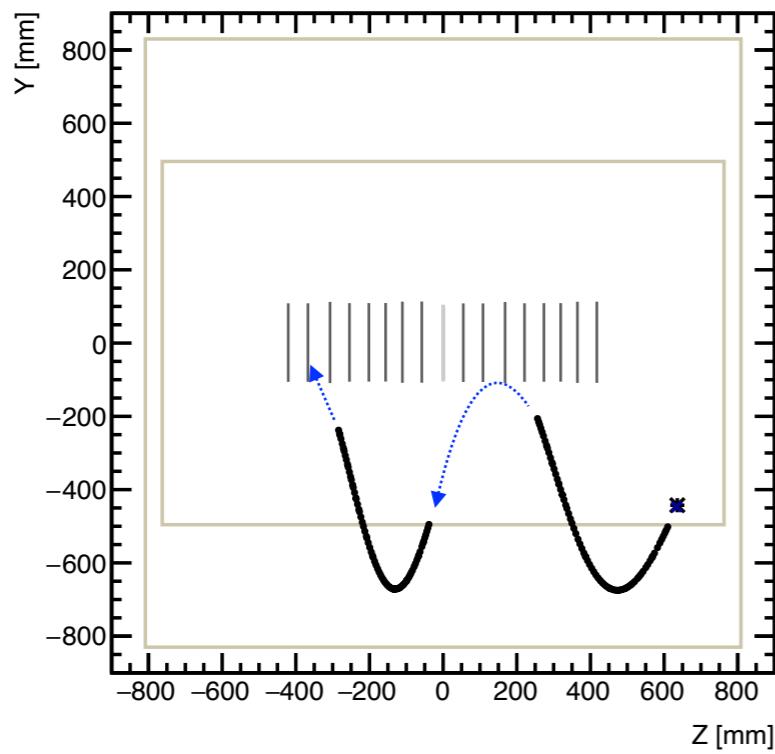
- ▶ Cosmic-ray muons is one of severest backgrounds in the COMET experiment.
- ▶ **Track-Direction ID** method was developed
 - to suppress **sneaking cosmic μ^+ BG**.
 - Demonstrated that **reverse μ^+ track** can be suppressed by an **order of magnitude**, with **signal e^- track retention efficiency of ~90%** (assuming 150 μm spatial resolution).
 - ✓ Note reduction power is sensitive to the CDC spatial resolution.

● Prospect

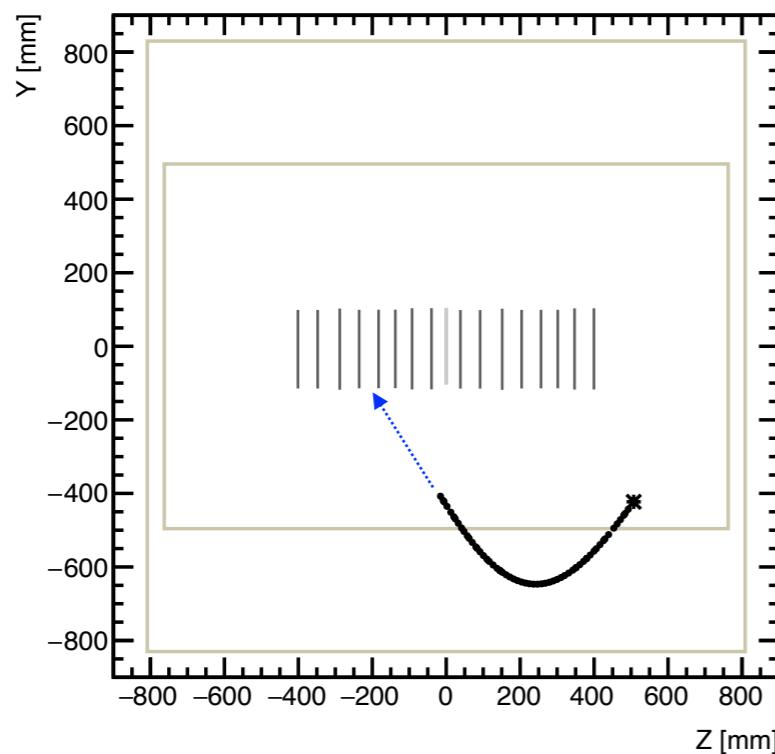
- A study is on going to improve reduction power by using machine learning with the noise hits.

Backup

Note for multiple-turn events



If a track has multiple turns, the direction will be easily identified by checking momentum attenuation in-between the turns.



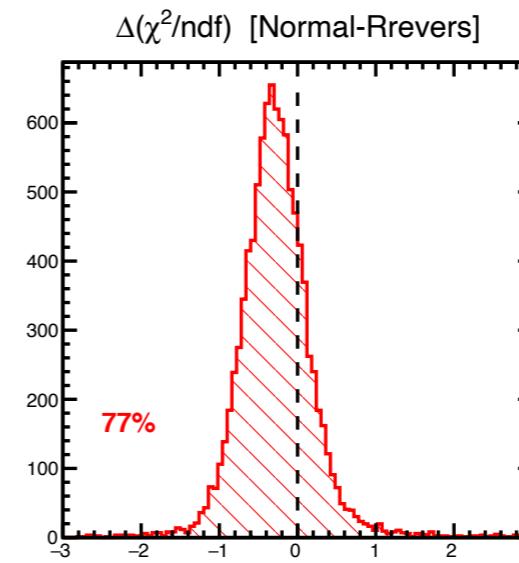
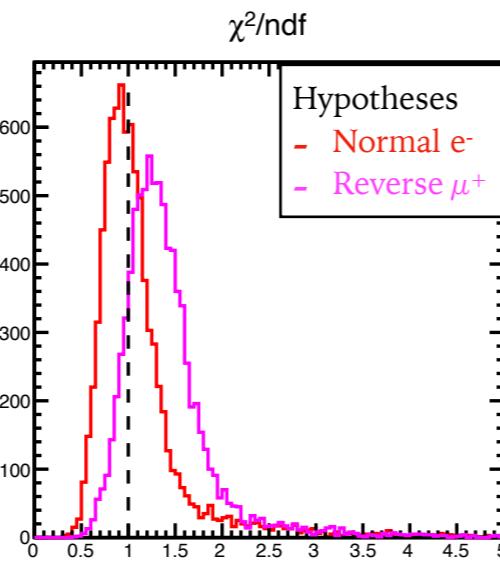
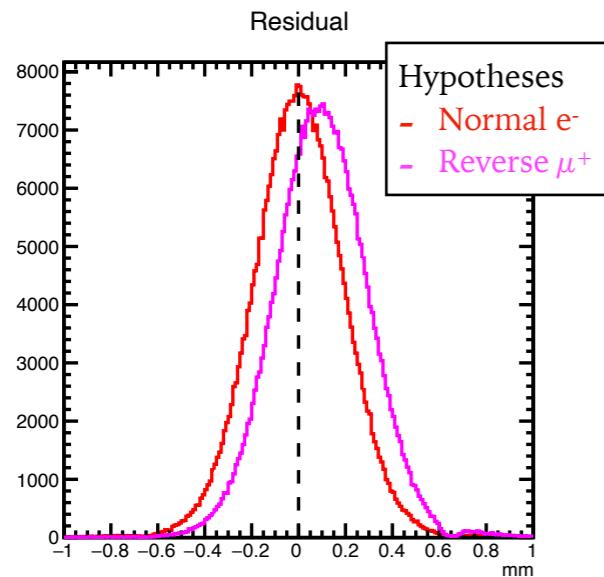
Therefore, we deal with only single-turn events in this study.

Spatial resolution = 200 μm

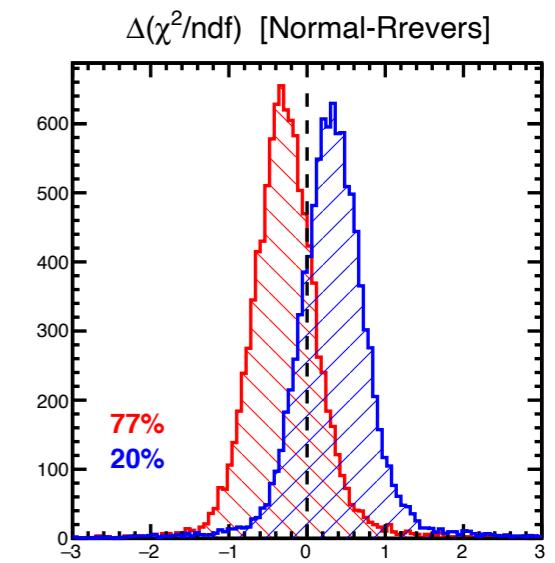
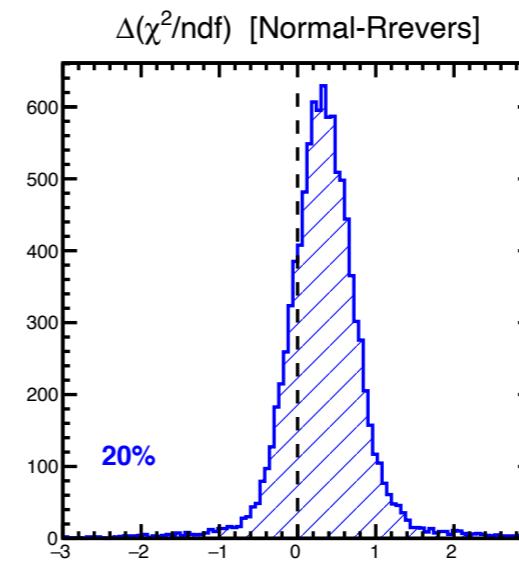
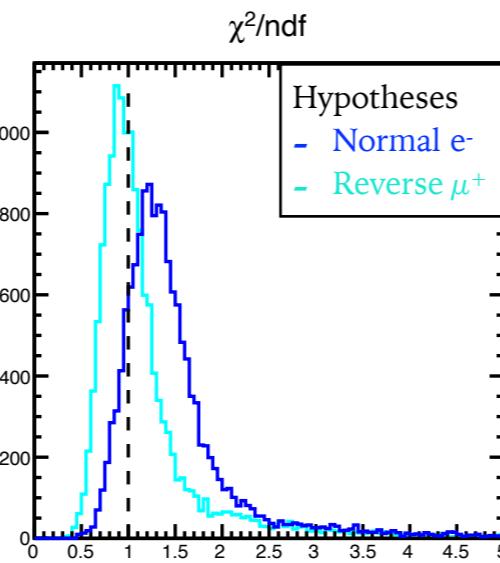
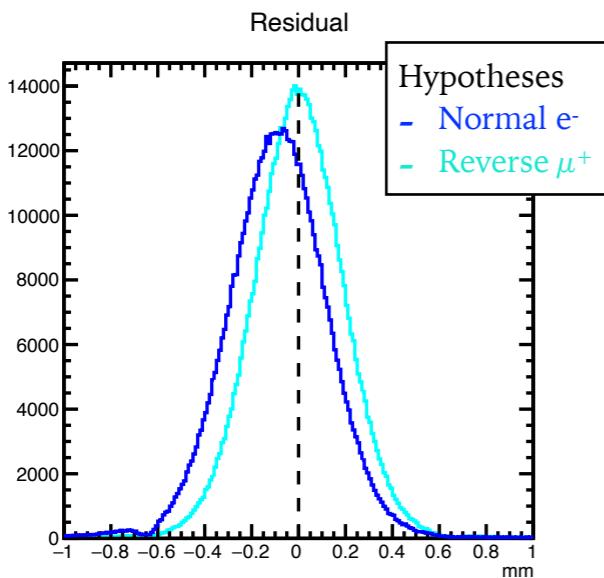
Compare 2 hypotheses

Normal-direction e^- vs Reverse-direction μ^+

Signal e^- MC sample



Reversed μ^+ MC sample

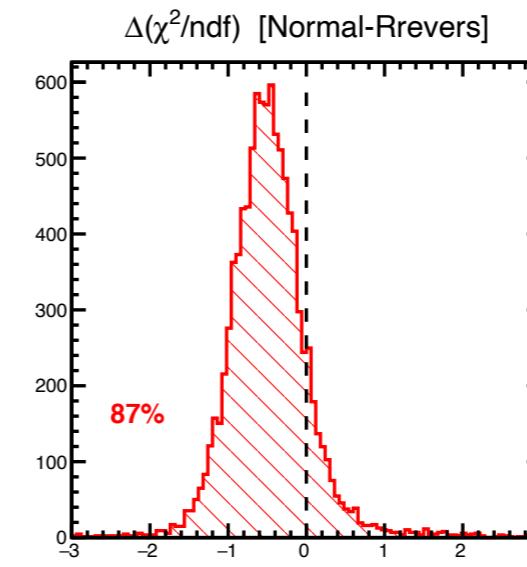
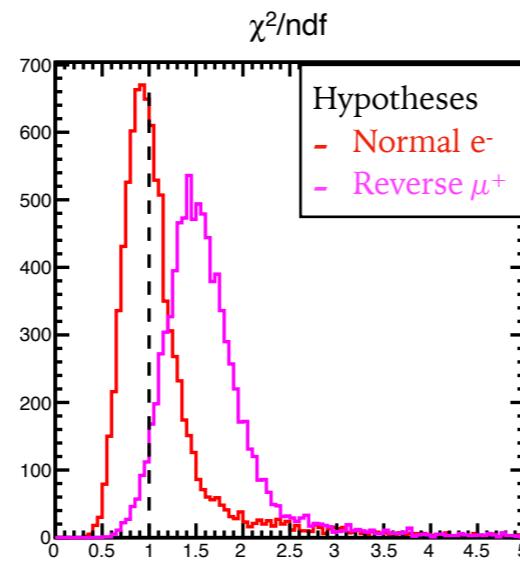
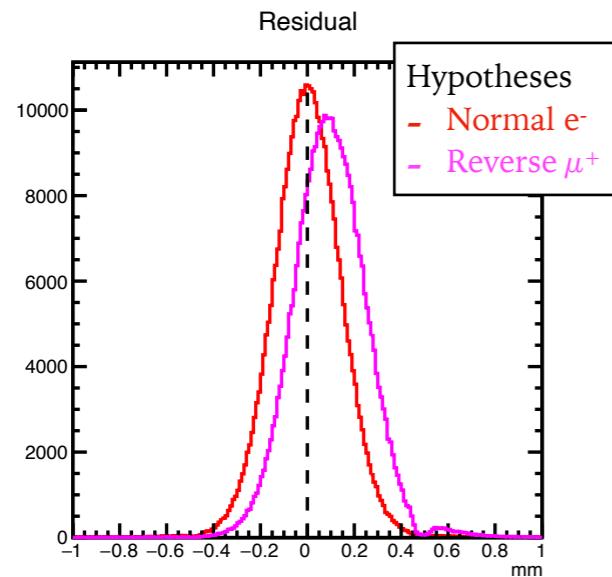


Spatial resolution = 150 μm

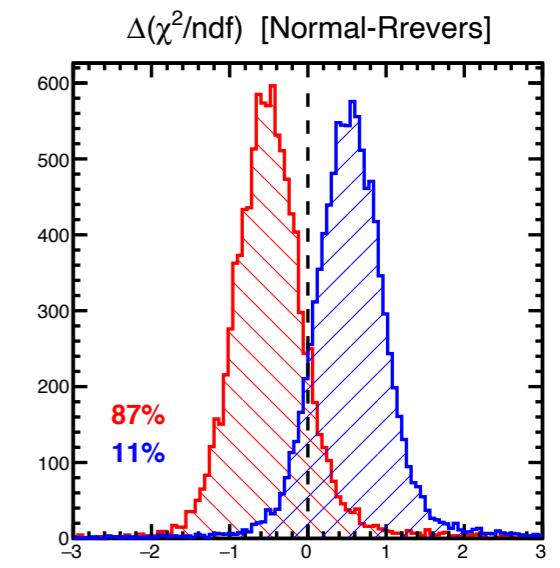
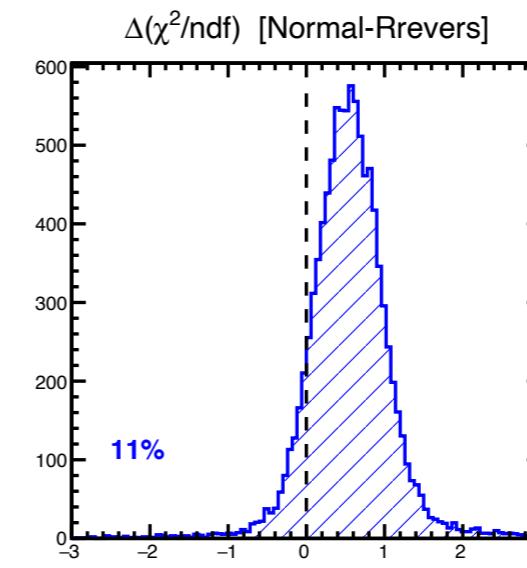
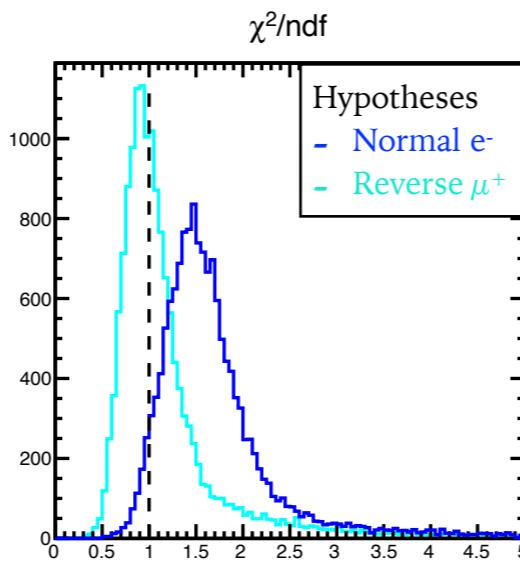
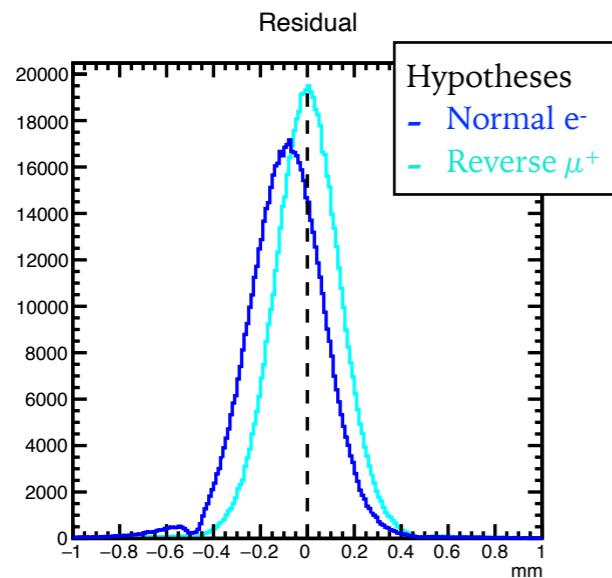
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Reversed μ^+ MC sample

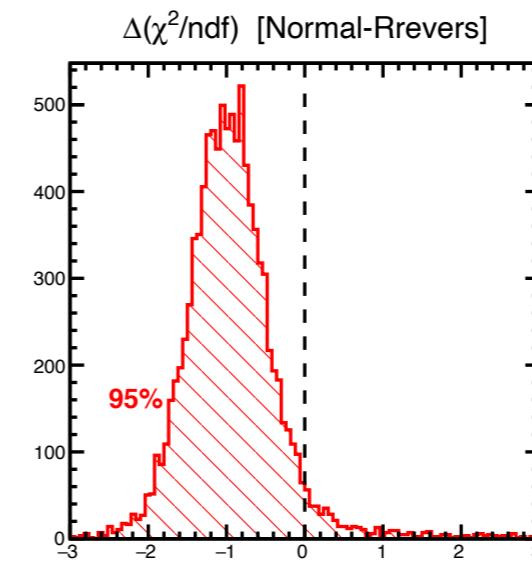
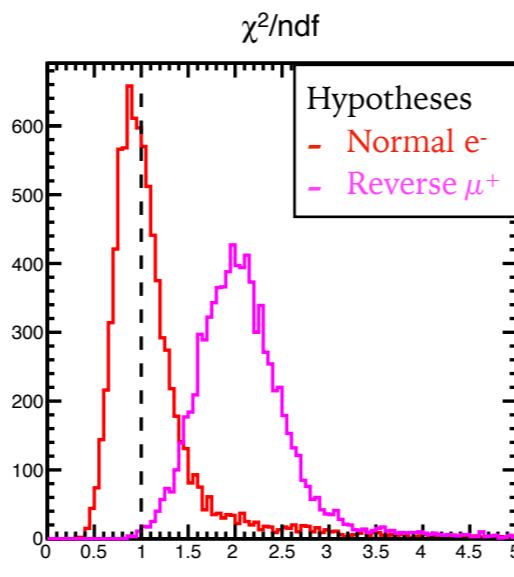
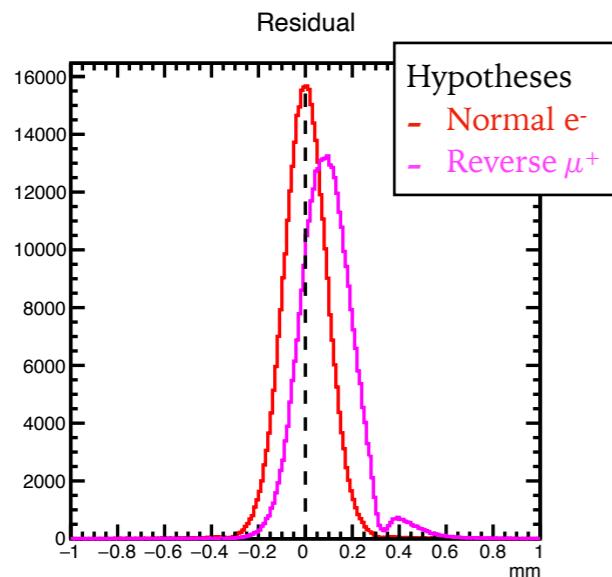


Spatial resolution = 100 μm

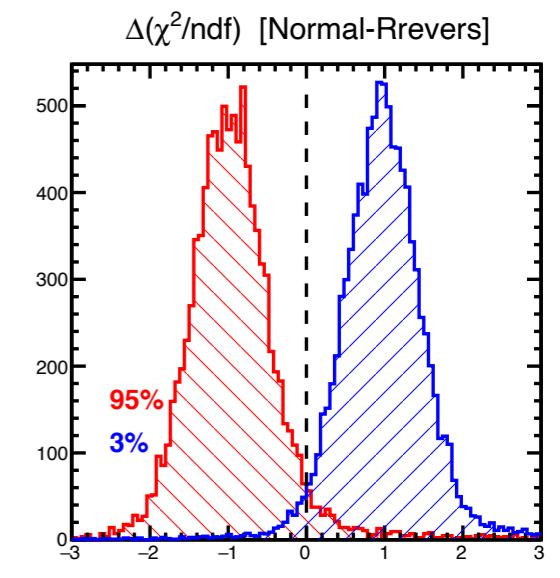
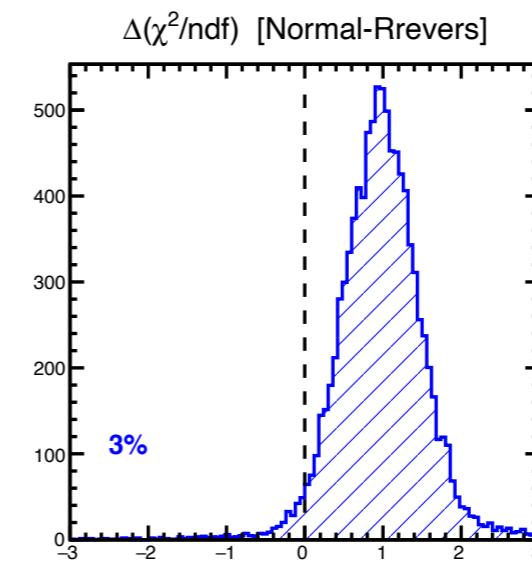
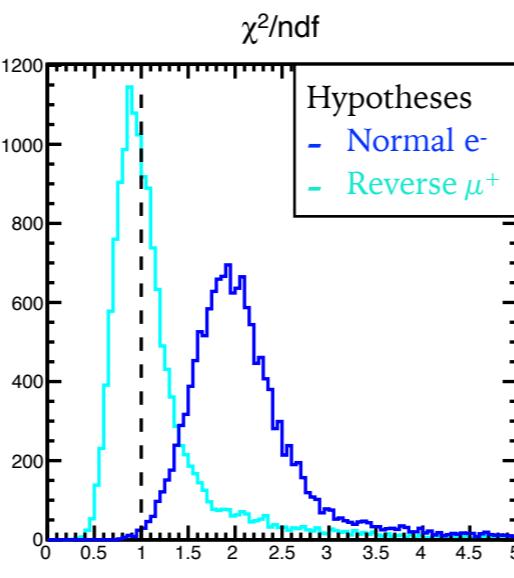
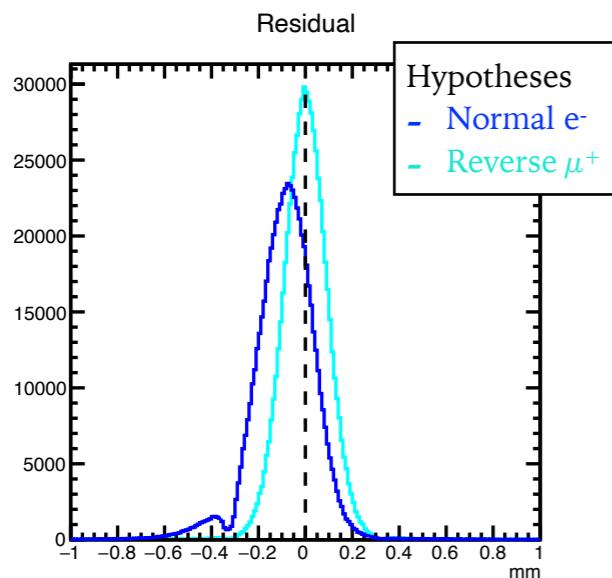
Compare 2 hypotheses

Normal-direction e^- vs Reverse-direction μ^+

Signal e^- MC sample

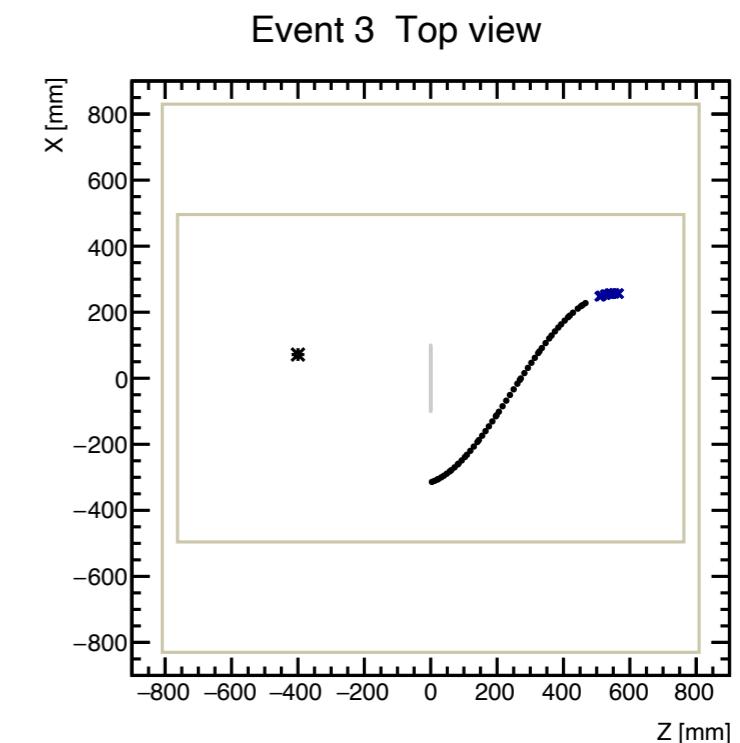
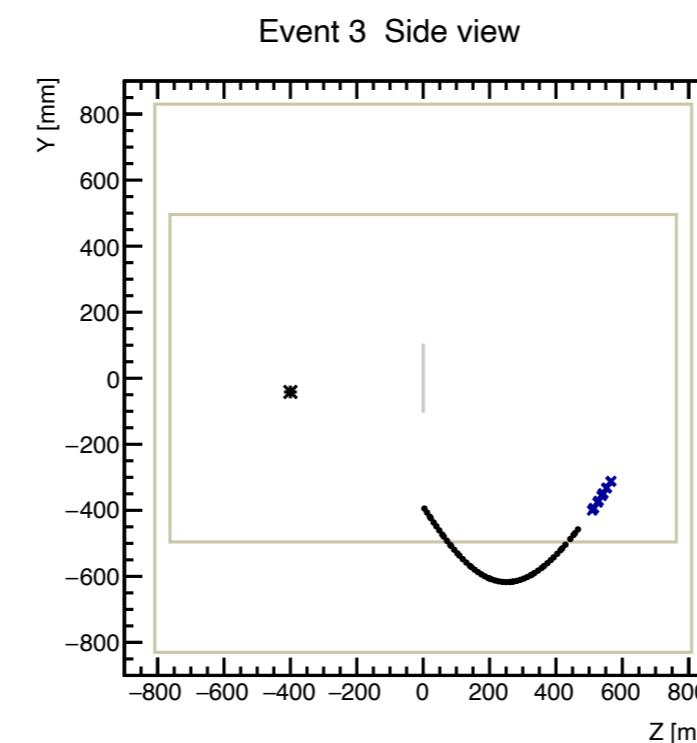
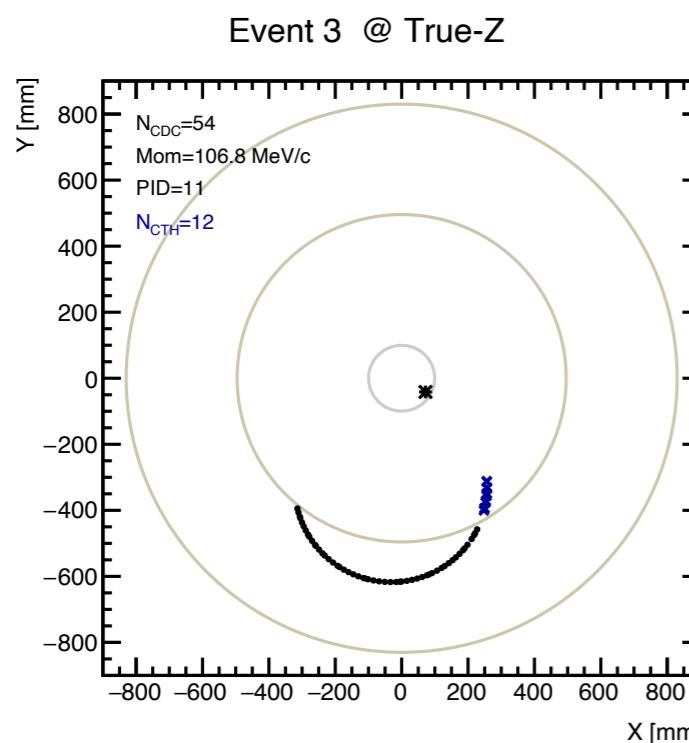
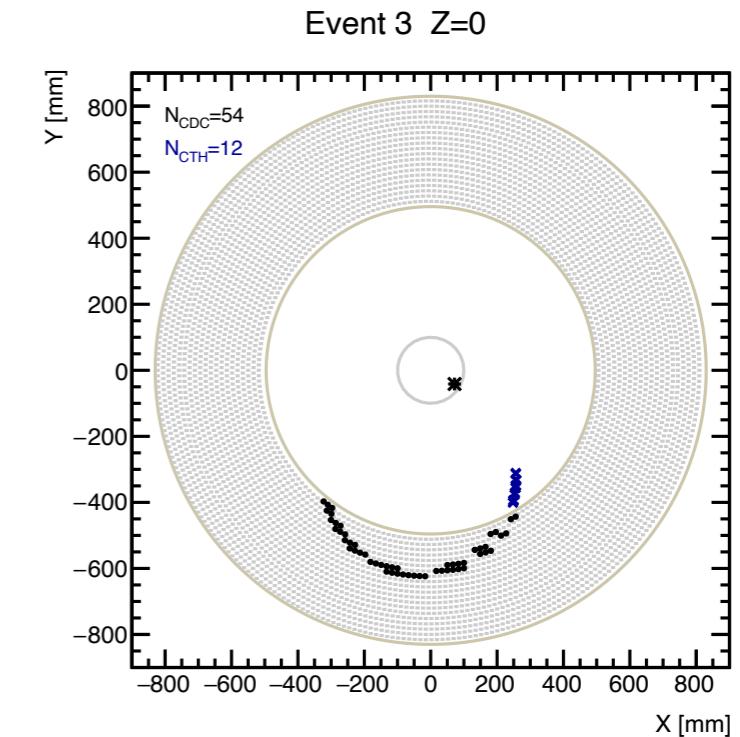
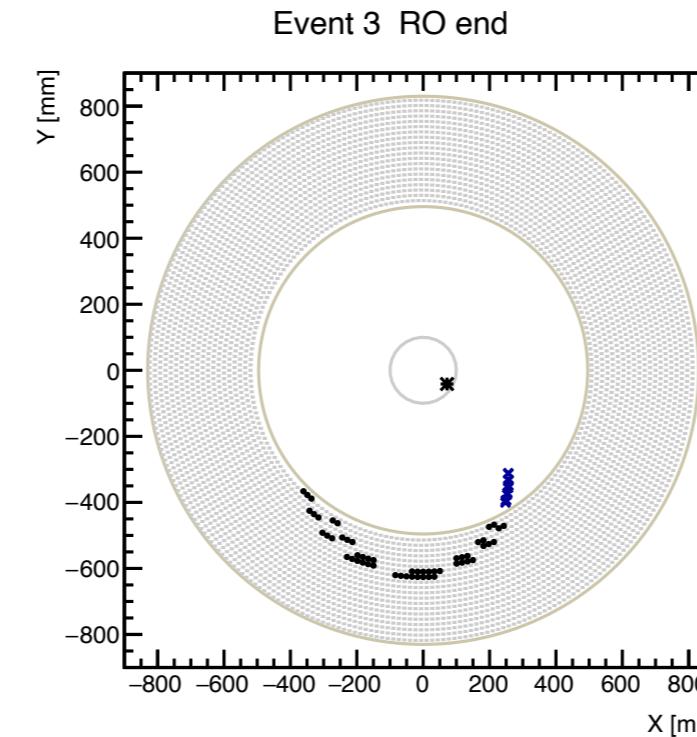
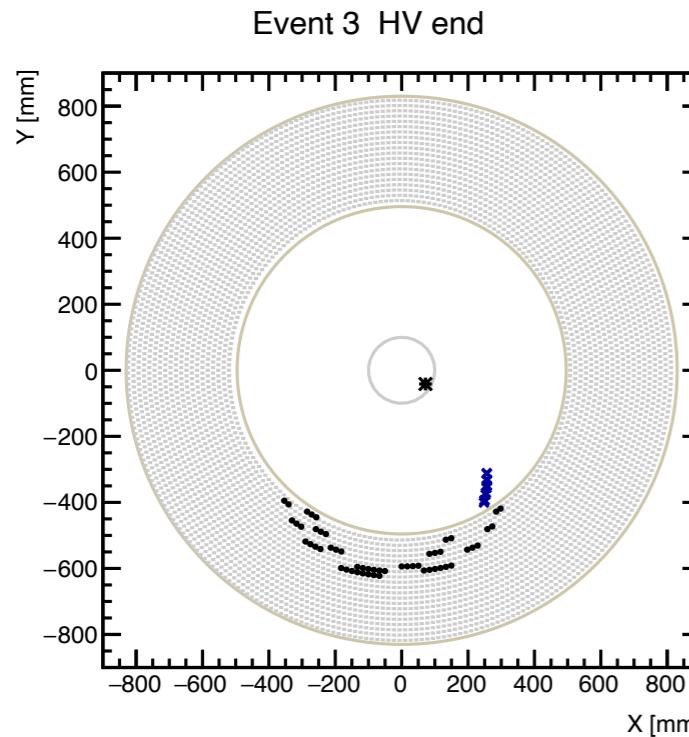


Reversed μ^+ MC sample



Event display example

Signal e- MC sample



How to discriminate reverse μ^+ ?

1. Cherenkov radiation

μ^+ with $\beta=0.7$ exceeds the Cherenkov threshold of Acrylic ($n=1.49$), $\beta_{th}=0.67$.

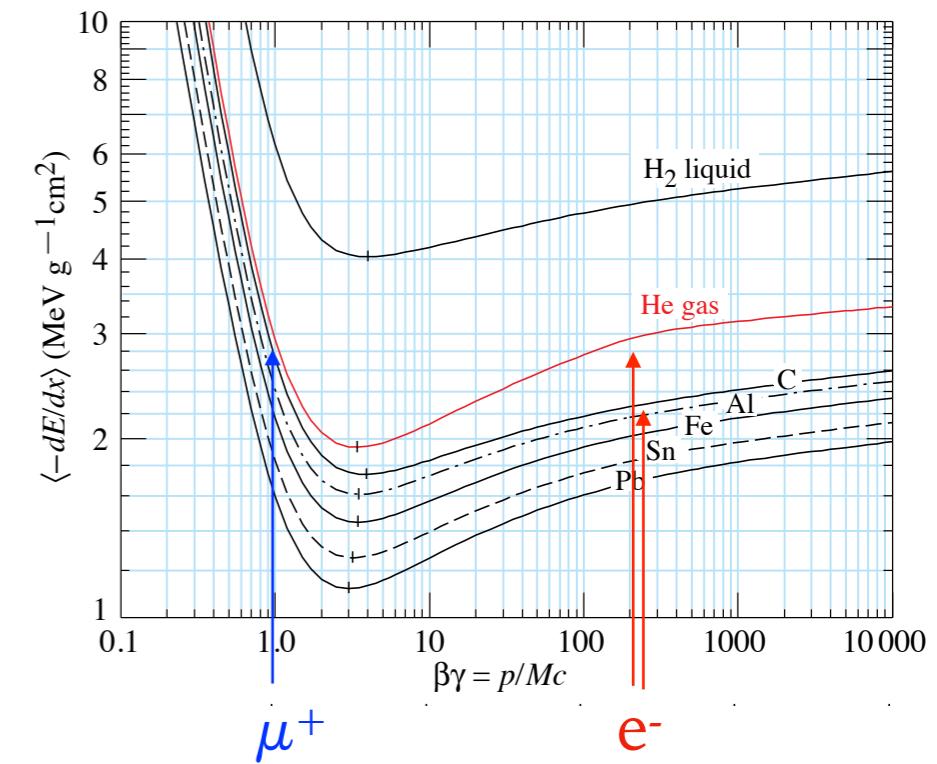
(Perhaps the number of photons may help to some extent.)

Radiator with refractive index less than 1.4 is the best.

How to discriminate reverse μ^+ ?

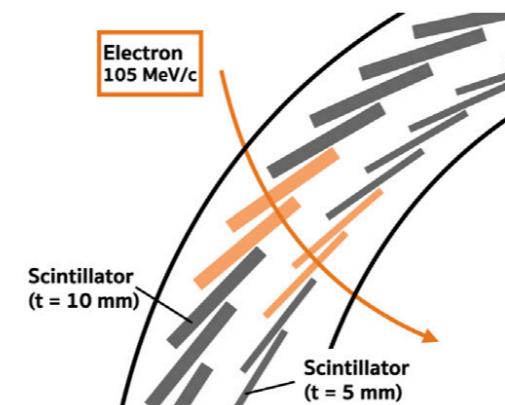
2. dE/dx

Unfortunately dE/dx in He gas for e^- ($\beta\gamma=200$) and μ^+ ($\beta\gamma=1$) are almost the same.
 → difficult to separate

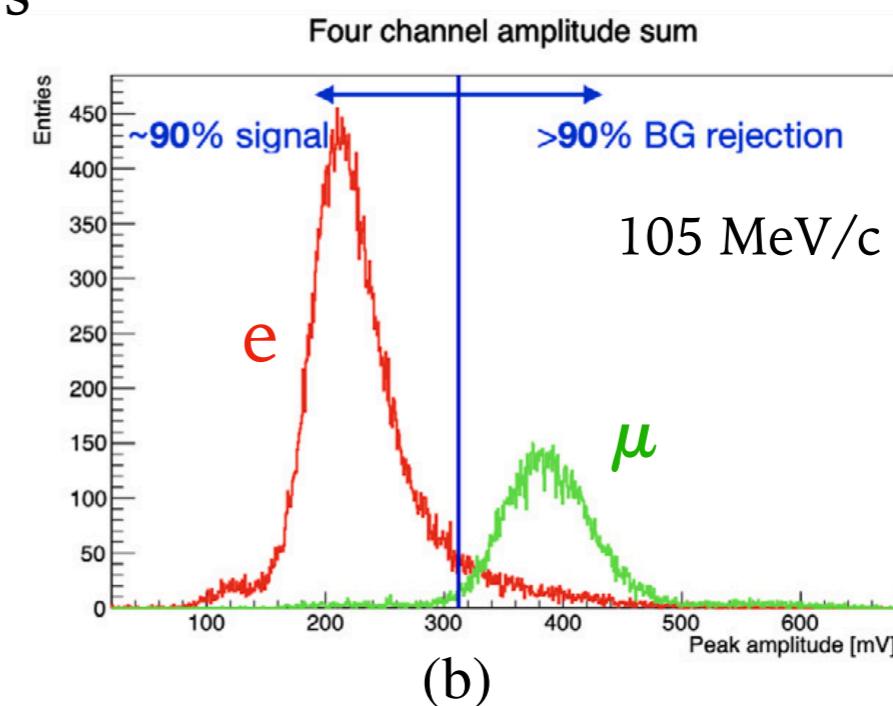


from PDG <http://pdg.lbl.gov/2019/reviews/rpp2018-rev-passage-particles-matter.pdf>

However, dE/dx in CTH Plastic Scintillators can have slight difference.



Y. Fujii et al. NIM A 1067, 169665 (2024)



Can we know track direction ?

We normally assume the track direction is from Target to CTH.

A. TOF miscorrection

In order to know correct drift time, time-0 is corrected according to TOF between CTH and each CDC hit.

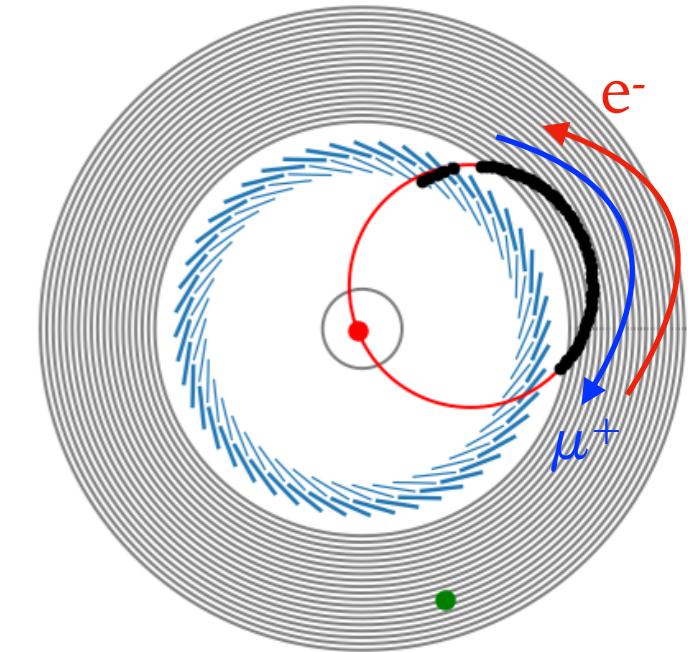
For **signal e^- ($\beta=1$)**, $-4 < \text{TOF} < 0 \text{ ns}$

For **reverse μ^+ ($\beta=0.7$)**, $0 < \text{TOF} < 5.7 \text{ ns}$

This causes miscorrection of **at longest 9.7 ns** for reverse μ^+ .

$$9.7 \text{ ns} \times 25 \mu\text{m/ns}^* = 240 \mu\text{m}$$

(comparable to spatial resolution of 150 μm)



* Drift velocity for He:C₄H₁₀ (90:10)
~ typically 25 $\mu\text{m/ns}$

Naive idea

This miscorrection will make the difference in χ^2 between normal & reverse direction hypotheses.

Can we know track direction ?

B. Material effect

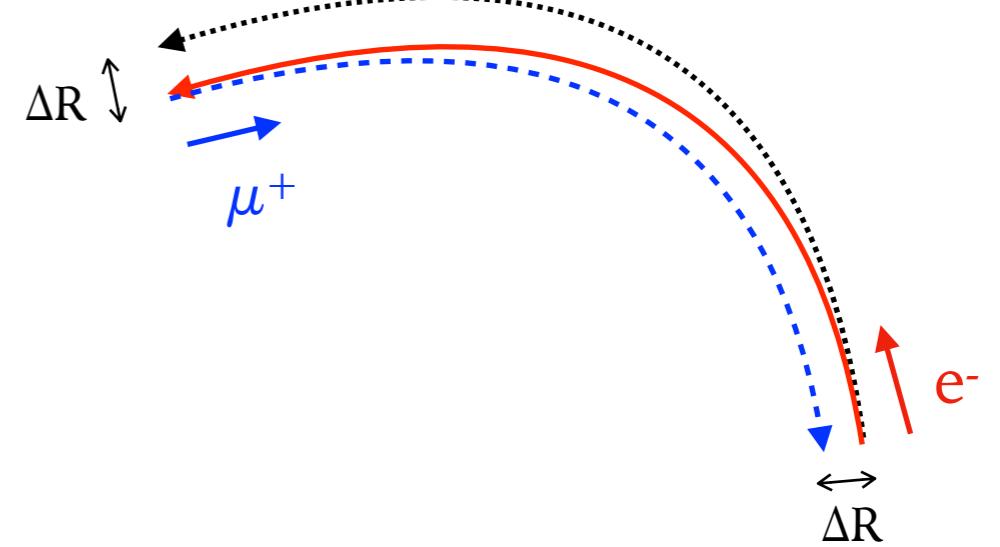
Energy loss is corrected in Kalman Filter

$dE \sim 1.5 \text{ keV}$ per cell

$\Delta E \sim 1.5 \text{ keV} \times 50 \text{ cells} = 75 \text{ keV}$ for track turn

$$\Delta R = \frac{\Delta p}{0.3B} = \frac{75 \text{ keV}}{0.3 \times 1 \text{ T}} = 250 \mu\text{m} \quad (\text{Rough estimation})$$

(comparable to spatial resolution of $150 \mu\text{m}$)



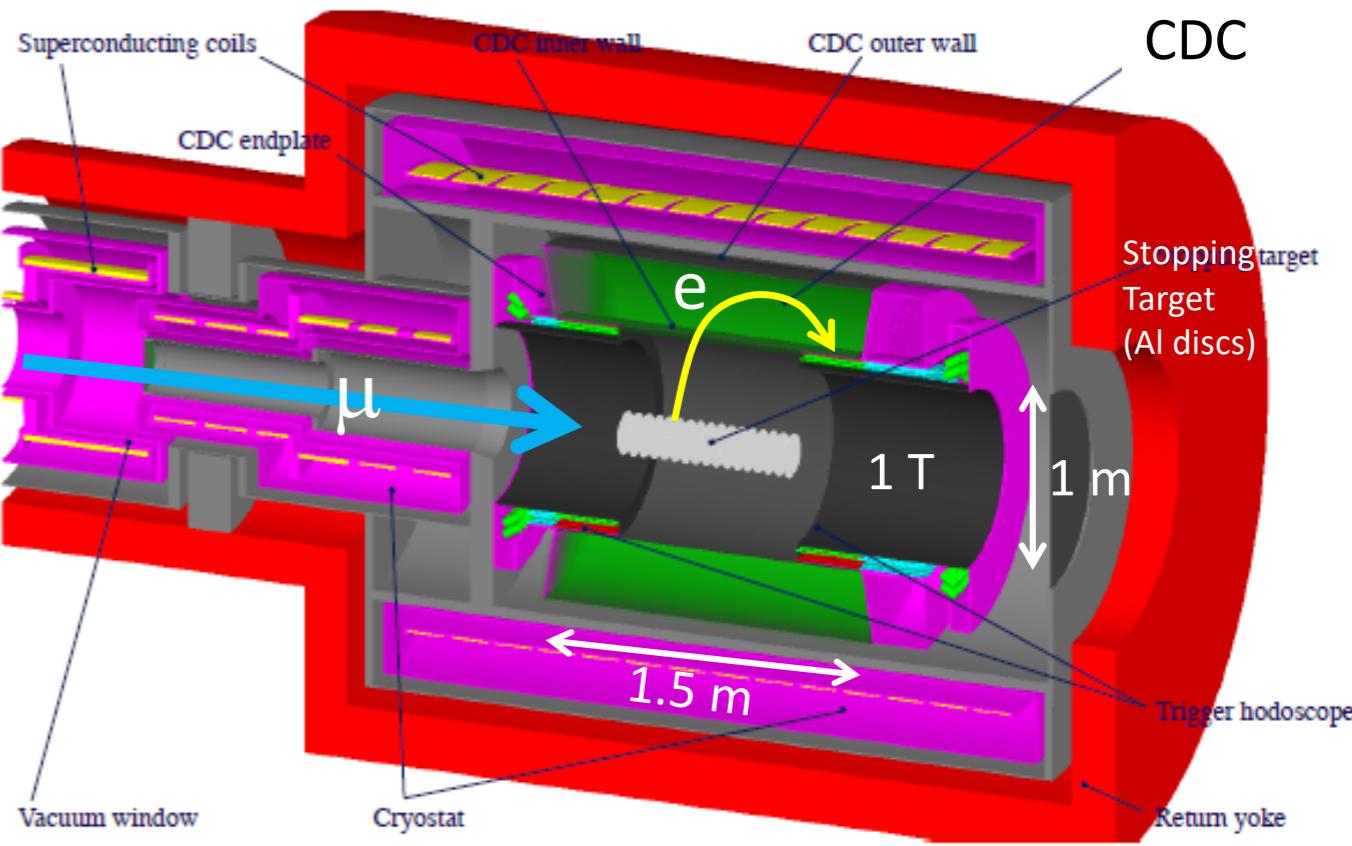
Energy loss correction also makes miscorrection
if the track direction is wrong.
→ worsen χ^2 !

COMET CDC

A. Sato et al., *NIM A* 1069, 169926 (2024)

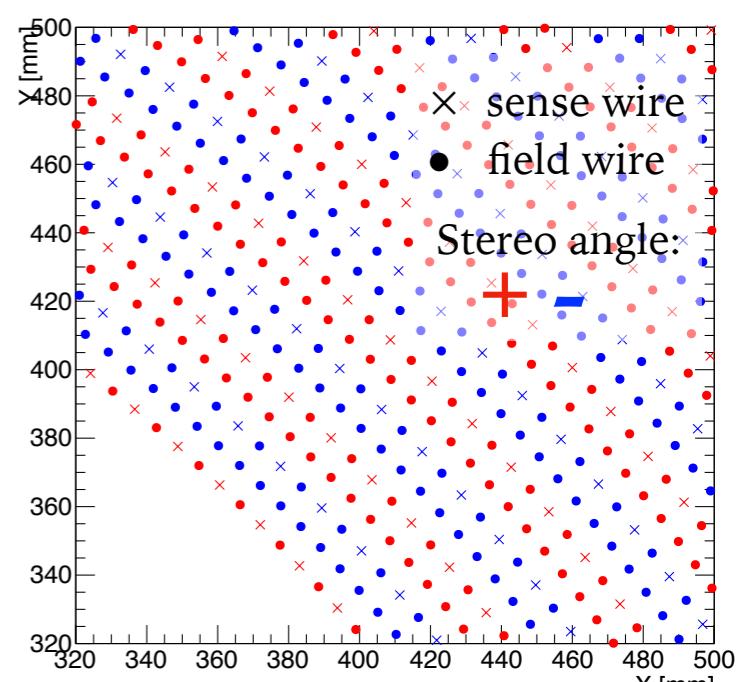
Design parameters of the CDC.

Inner wall	Material	CFRP, Al		
	Length	1495.5 mm		
	Radius	495.95–496.5 mm		
	Thickness	CFRP: 0.5 mm, Al: 0.05 mm		
Outer wall	Material	CFRP, Al	CDC	
	Length	1577.3 mm	Stopping ^{target}	
	Radius	829.9–835.0 mm	Target (Al discs)	
	Thickness	CFRP: 5.0 mm, Al: 0.1 mm		
Number of sense layers	20 (including 2 guard layers)			
Sense wire	Material	Au-plated W		
	Diameter	25 μ m		
	Number of wires	4986		
	Tension	50 g		
	Stereo angle	$\pm(64\text{--}75)$ mrad		
Field wire	Material	Al alloy (A5056)		
	Diameter	126 μ m		
	Number of wires	14 562		
	Tension	80 g		
	Stereo angle	$\pm(64\text{--}75)$ mrad		
Gas	Mixture	He:i-C ₄ H ₁₀ (90:10)		
	Volume	2084 L		



Features

- ▶ Low mass gas = He:iC₄H₁₀ (90:10) ➔ p resolution
- ▶ Alternative all stereo wire, ± 4 deg ➔ z resolution
- ▶ Large inner bore, $\phi 1$ m @ 1 T ➔ suppress DIO hits



20 layers in total

Track Fitting using GENFIT

We implemented GENFIT into the COMET software framework (ICEDUST).

GENFIT

- Generic Track Fitting Framework
- Experiment-independent, modular packages
- Open source C++ code [GenFit in GitHub](#)
- Originally developed in PandaROOT at TUM [Hoppner et al. NIM A 620, 518 \(2010\)](#)
- Widely used in many experiments, e.g. Belle-II, SHiP, FOPI etc.
- Suitable for low-energy experiments [Bilka et al., arXiv:1902.04405](#)
- Fitter Options:
 - Kalman Filter, [Deterministic Annealing Filter](#) etc.

Works well !!

