

Muon Conversion Experiments: COMET/Mu2e and future

MyeongJae Lee

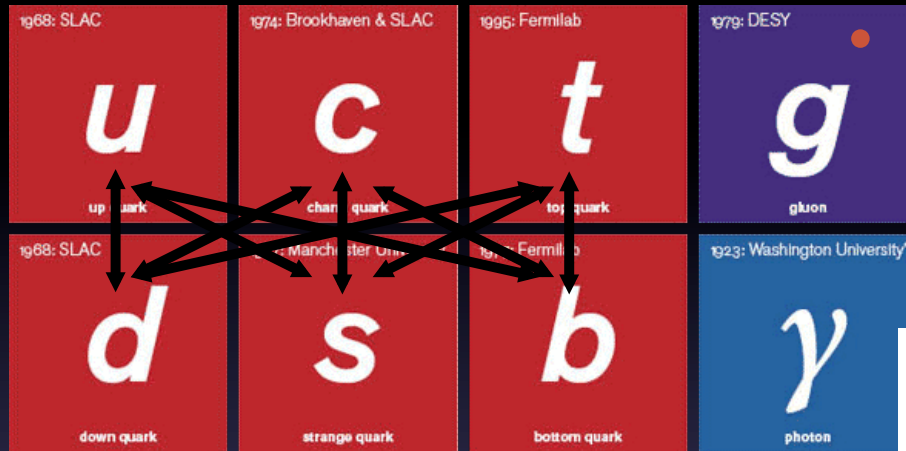
(Institute for Basic Science, Korea)

Oct 14, 2015

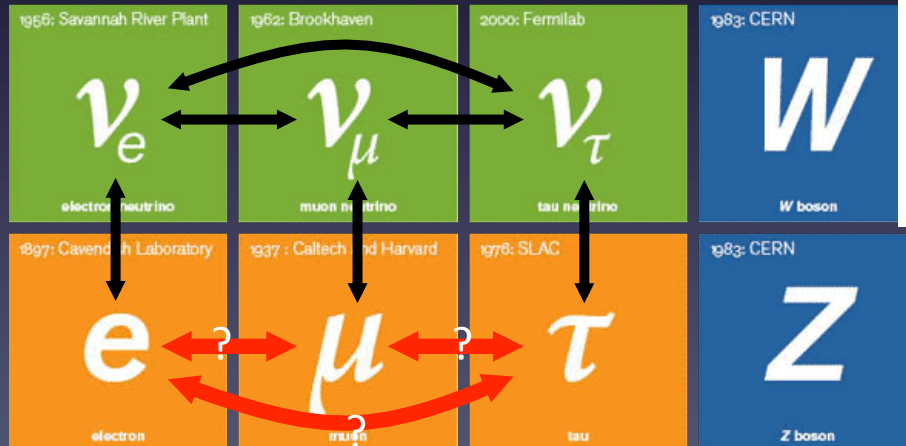
HINT 2015

Lepton Flavor Violation (LFV)

QUARKS

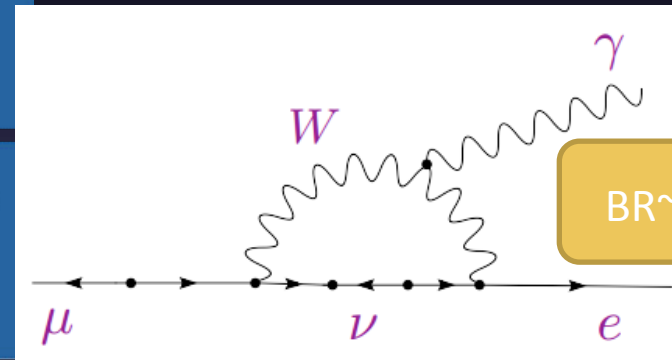


LEPTONS



Charged Lepton mixing
NOT observed

— Very small possibility in SM



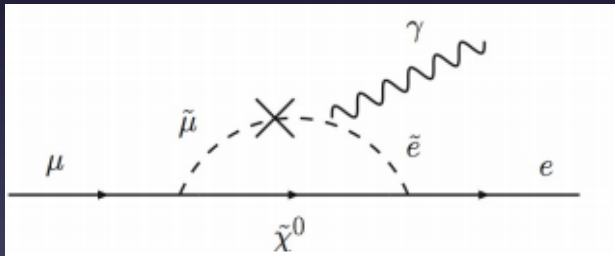
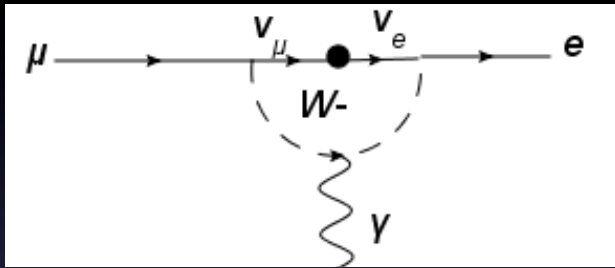
— **LFV = New physics** in BSM

GAUGE
BOSONS

Muon LFV

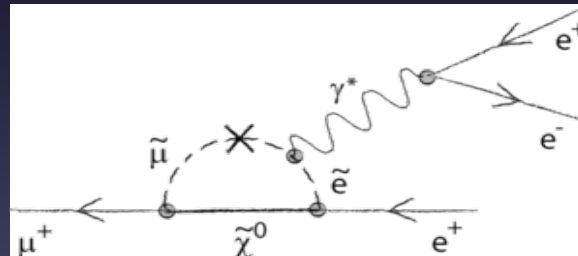
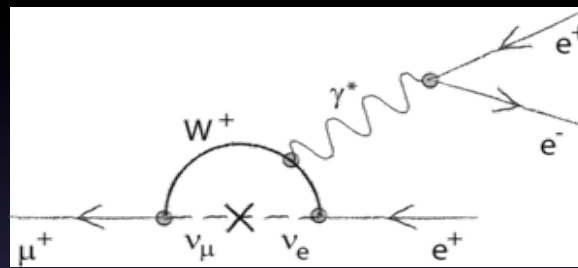
Note : $\mu \rightarrow e \nu_e \nu_\mu$

$\mu \rightarrow e \gamma$



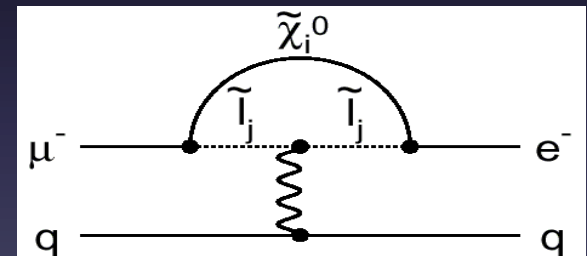
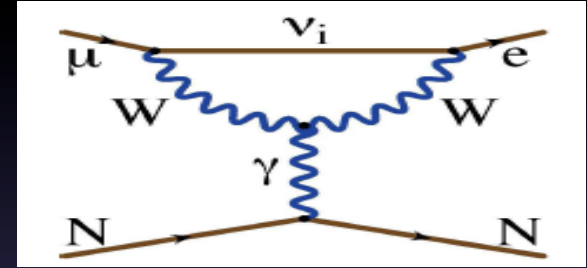
- MEG (PSI)
- MEG upgrade (planned)

$\mu \rightarrow 3e$



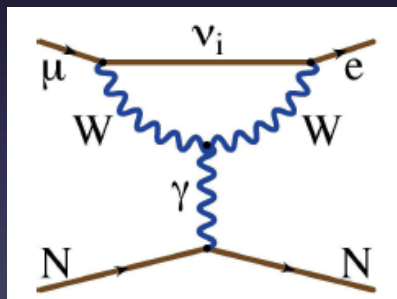
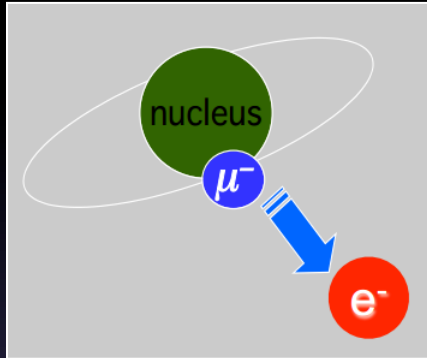
- Mu3e (PSI, planned)

$\mu N \rightarrow e N$



- Mu2e (FNAL)
- **COMET** (J-PARC)

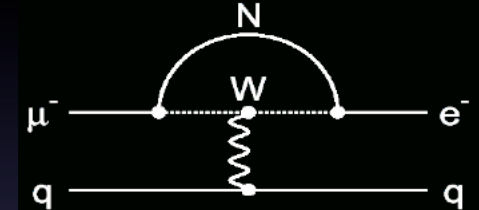
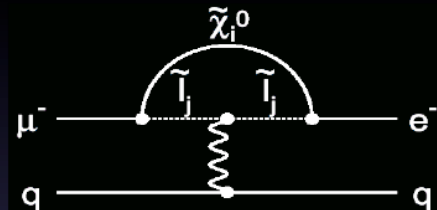
Muon conversion experiment



Standard model Muon conversion : $R \approx 10^{-54}$

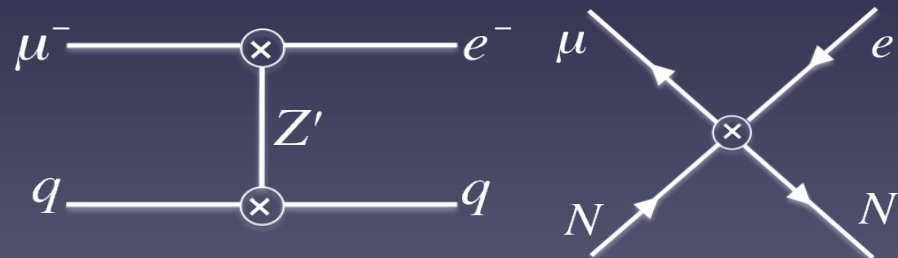
$$R_{\mu e} = \frac{\text{Muon to Electron conversion rate}}{\text{Ordinary muon capture rate}}$$

Beyond Standard Model



Supersymmetry : $R \leq 10^{-13}$ Heavy neutrino

Only COMET/Mu2e sensitive

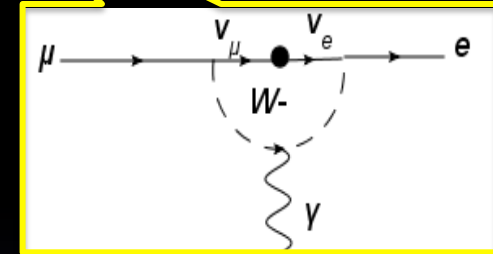
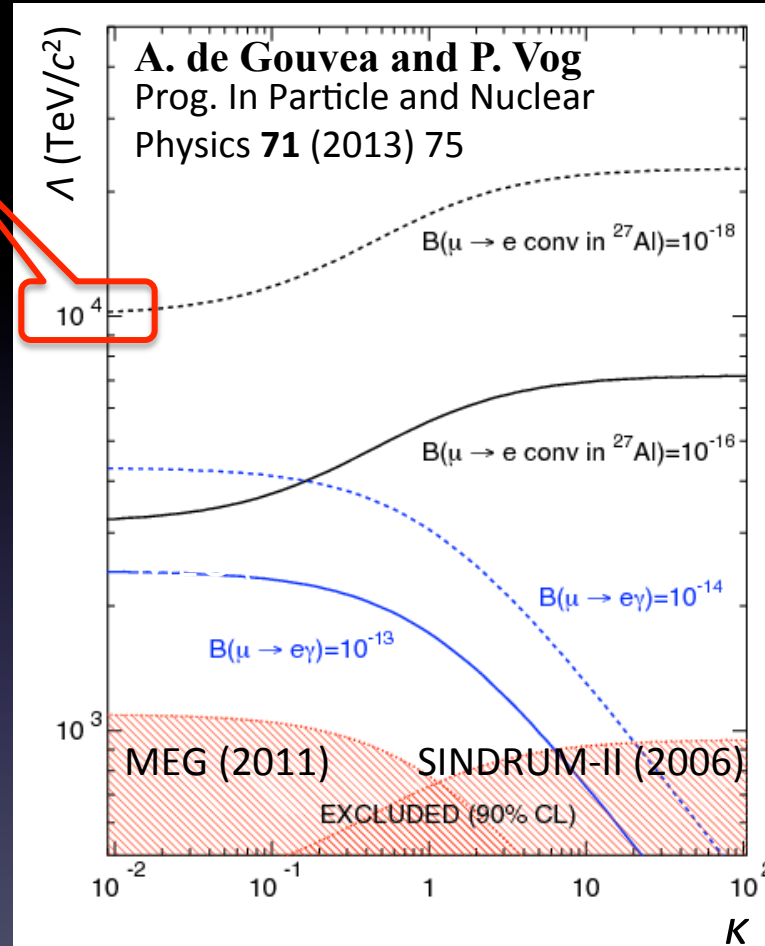
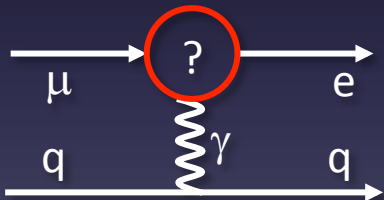


$\mu N \rightarrow e N$ vs $\mu \rightarrow e \gamma$

Probing $O(10^4)$ TeV mass scale,

→ Much higher energy scale than LHC

Dipole Interaction

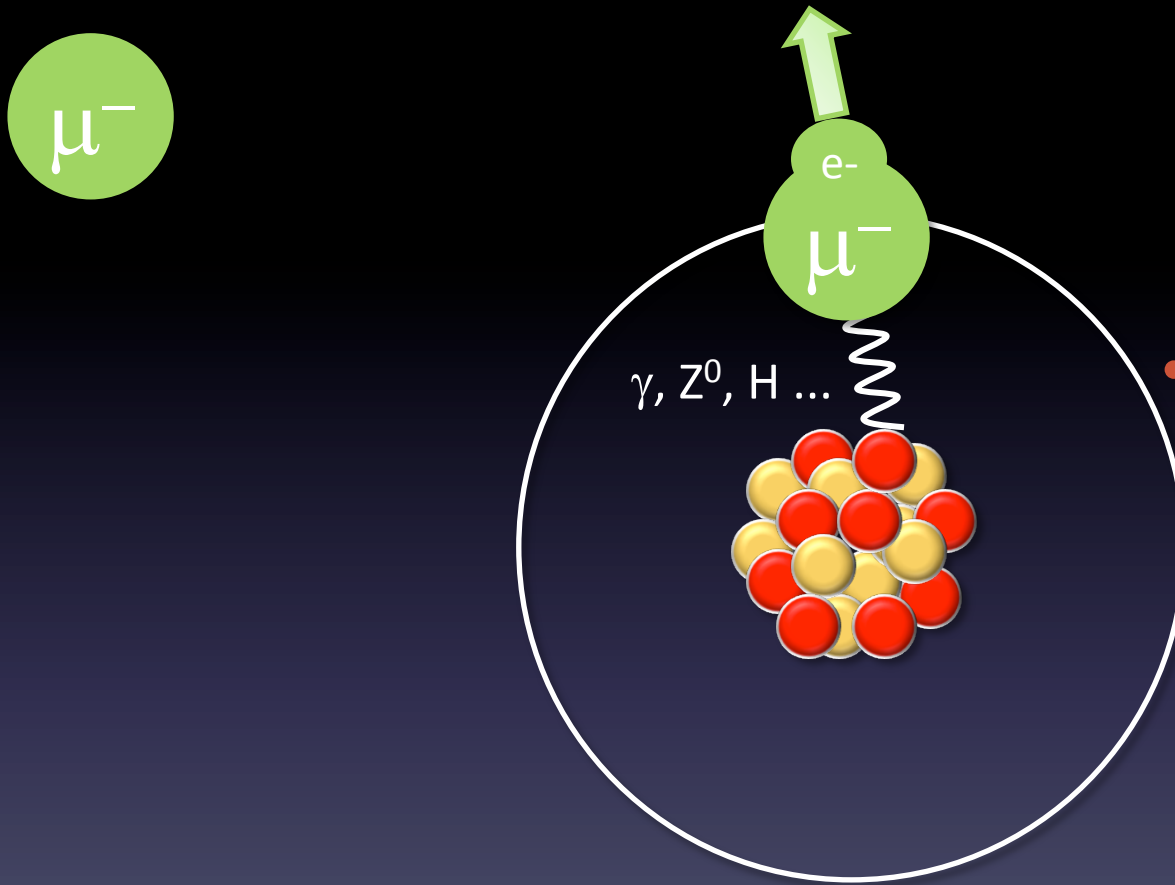


Contact interaction



$$L_{CLFV} = \frac{m_\mu}{(1 + \kappa)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(1 + \kappa)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L \left(\sum_{q=u,d} \bar{q}_L \gamma_\mu q_L \right)$$

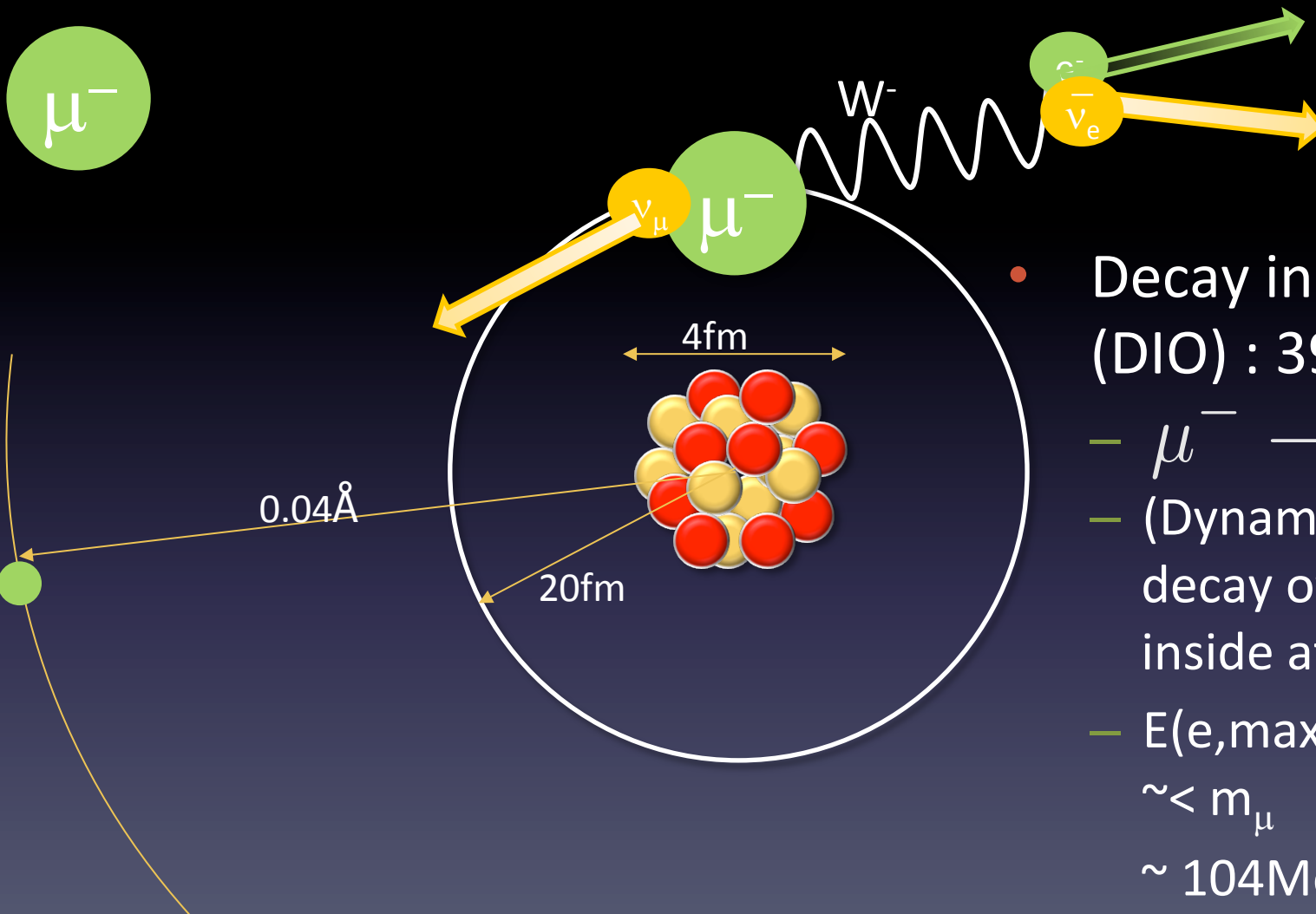
Stopped μ^- in matter: Conversion



• Muon conversion

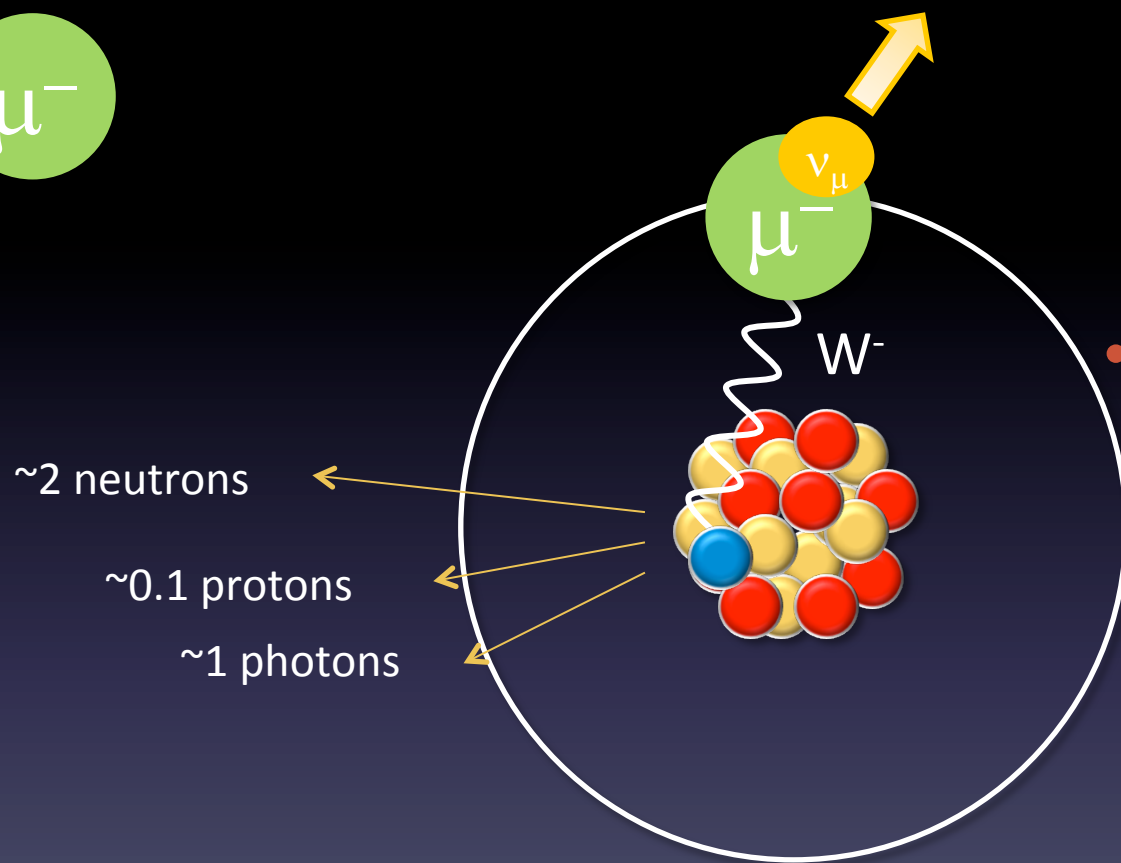
- $\mu^- + N \rightarrow e^- + N$
- Muon conversion coherently with nucleus

Stopped μ^- in matter: DIO



- Decay in orbit (DIO) : 39% (Al)
 - $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$
 - (Dynamically) free decay of muon inside atom
 - $E(e, \text{max}, \text{Al}) \sim < m_\mu \sim 104 \text{ MeV}$ (peak at 50 MeV)

Stopped μ^- in matter: Capture

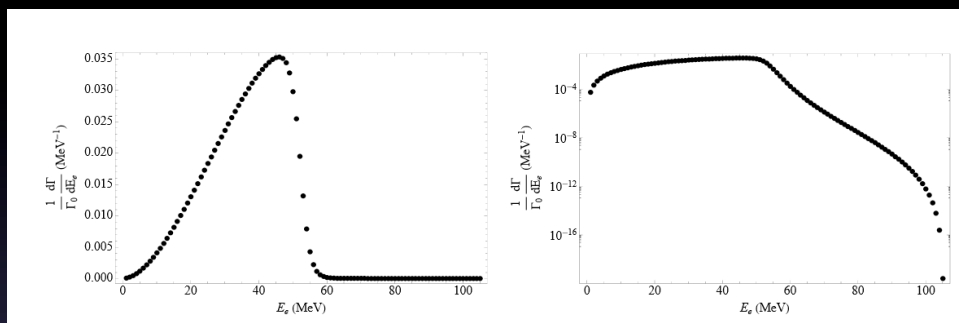


• Muon capture :
61% (Al)

- $\mu^- + p \rightarrow \nu_\mu + n$
- Muon decay coherently with nucleus
- Source of background hits

Backgrounds

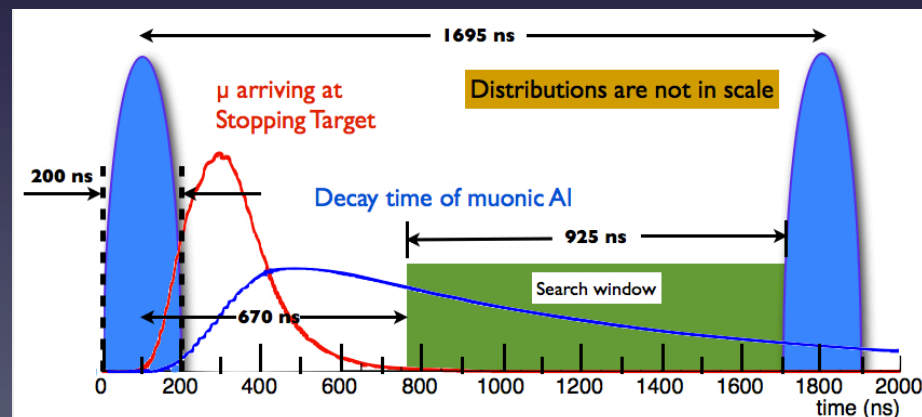
- Decay-In-Orbit (DIO) : dominant BG
 - Not always BG but one way to calibrate detector



- Prompt background (to beam arrival)

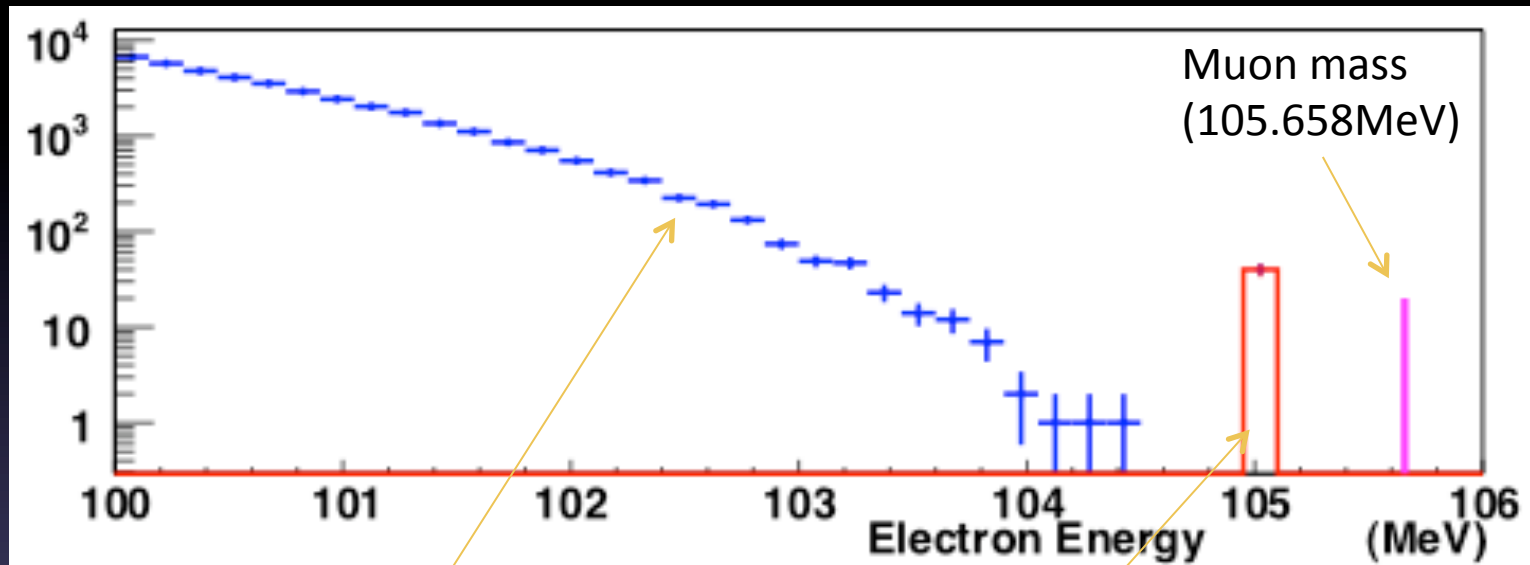
- By beam $\pi/n/e$, muon/pion decay in flight
- Ex) $\pi N \rightarrow \gamma X$; $\gamma \rightarrow e^+e^-$, $E(e^-) \sim 105 \text{ MeV}$ mimics signal
- “Pulsed beam” with “high extinction” required

(Mu2e beam spill profile)

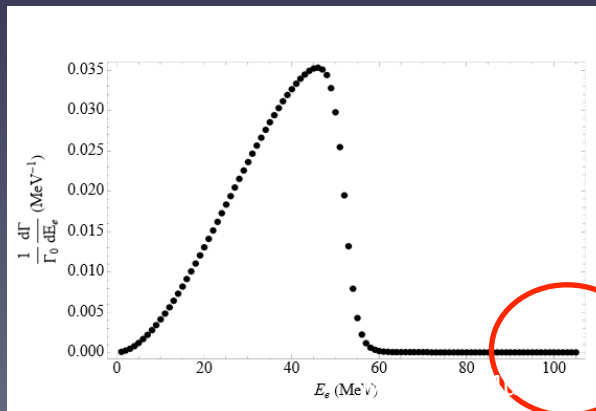


- 10^{-10} extinction required, achieved $10^{-10} - 10^{-12}$ at COMET

Muon Conversion Signal (Ideal detector)

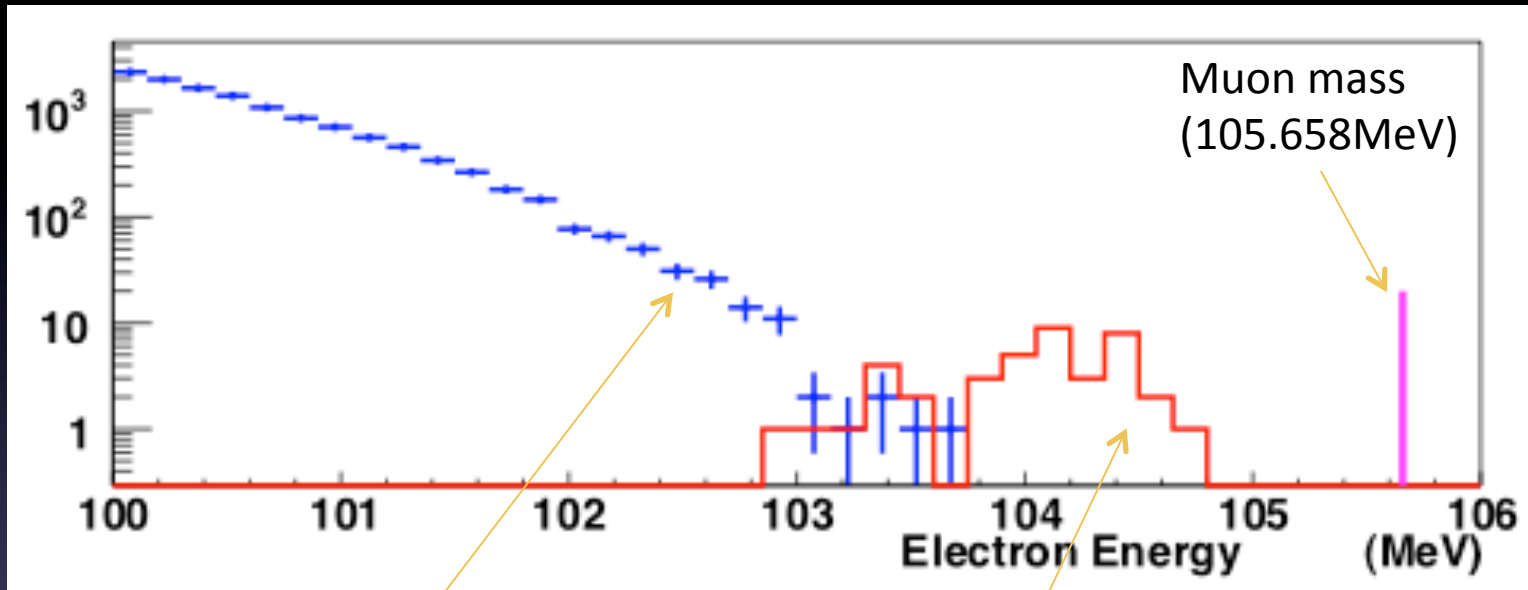


Muon Decay-In-Orbit

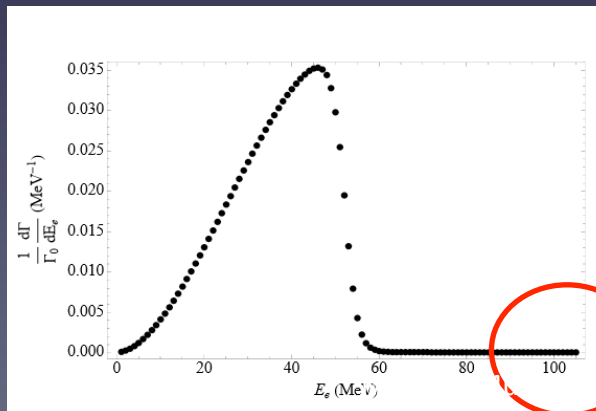


Signal : ~105 MeV
monoenergetic electron
 $= m_\mu - E_{\text{rec}} - E_B$

Muon Conversion Signal (Real detector)

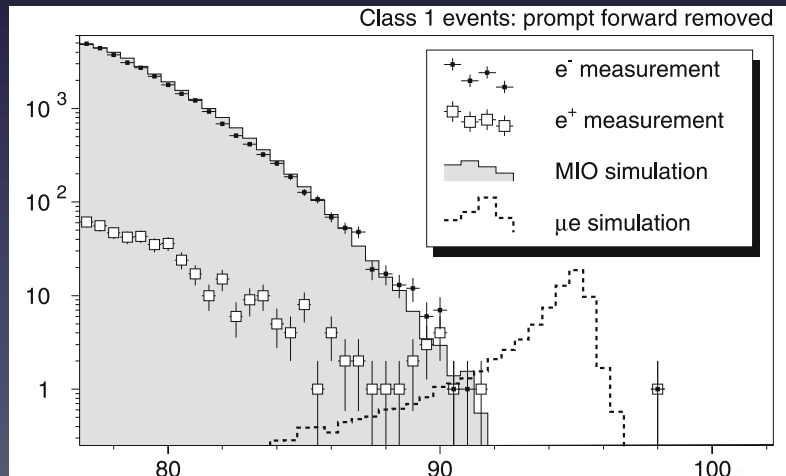
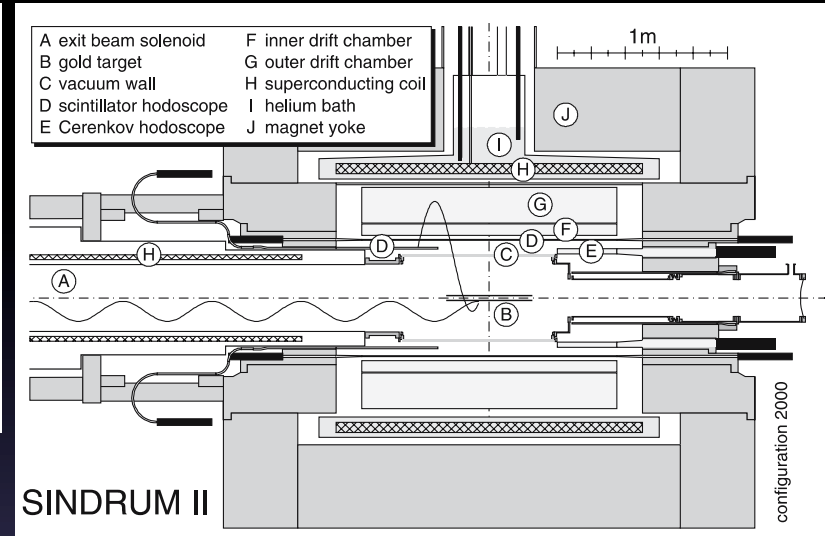
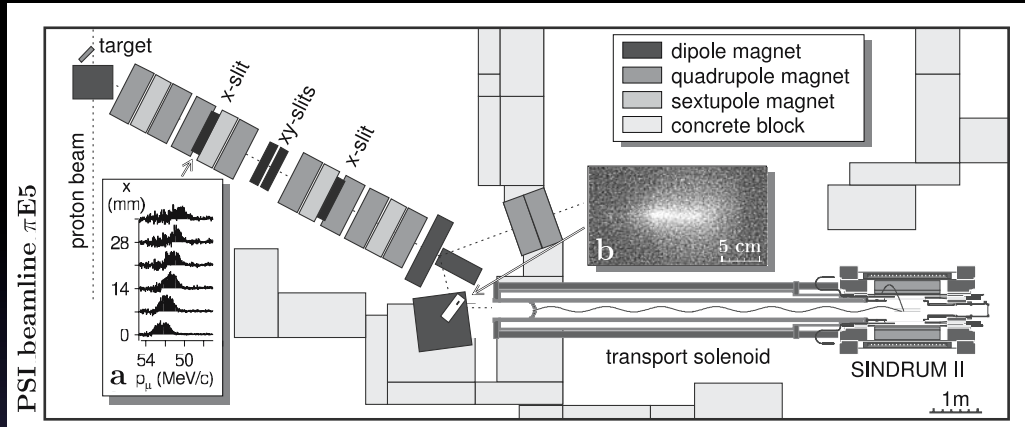


Muon Decay-In-Orbit



Signal : ~105 MeV
monoenergetic electron
 $= m_\mu - E_{\text{rec}} - E_B$

A precursor : SINDRUM-II (@PSI)



$$B_{\mu e}^{\text{Au}} < 7 \times 10^{-13} \quad 90\% \text{ C.L.}$$

Eur. Phys. J. C 47, 337–346 (2006)
Digital Object Identifier (DOI) 10.1140/epjc/s2006-02582-x

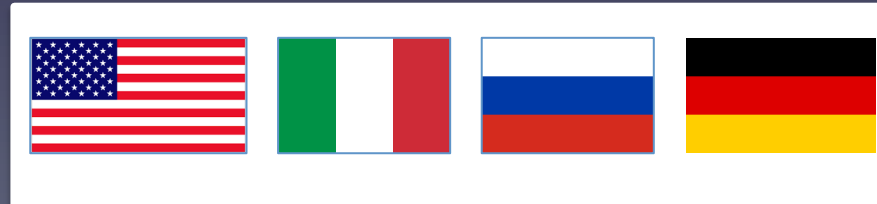
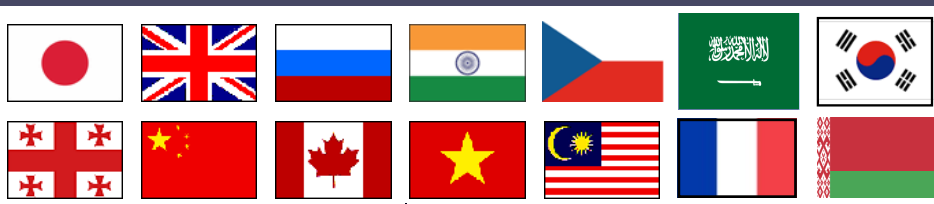
COMET (J-PARC) / Mu2e (FNAL)



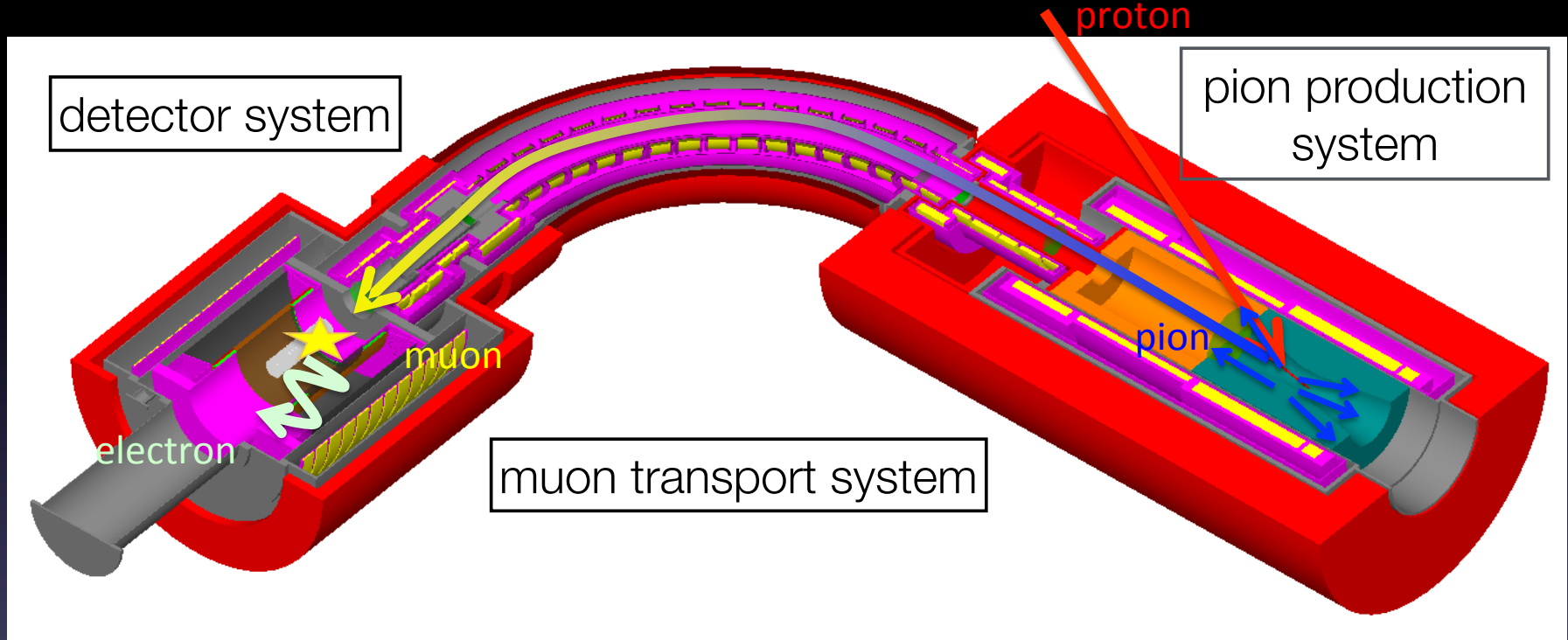
177 collaborators,
32 institutes
(as of July 2015)



193 collaborators
33 institutions
(as of June 2015)



COMET Phase-I experiment



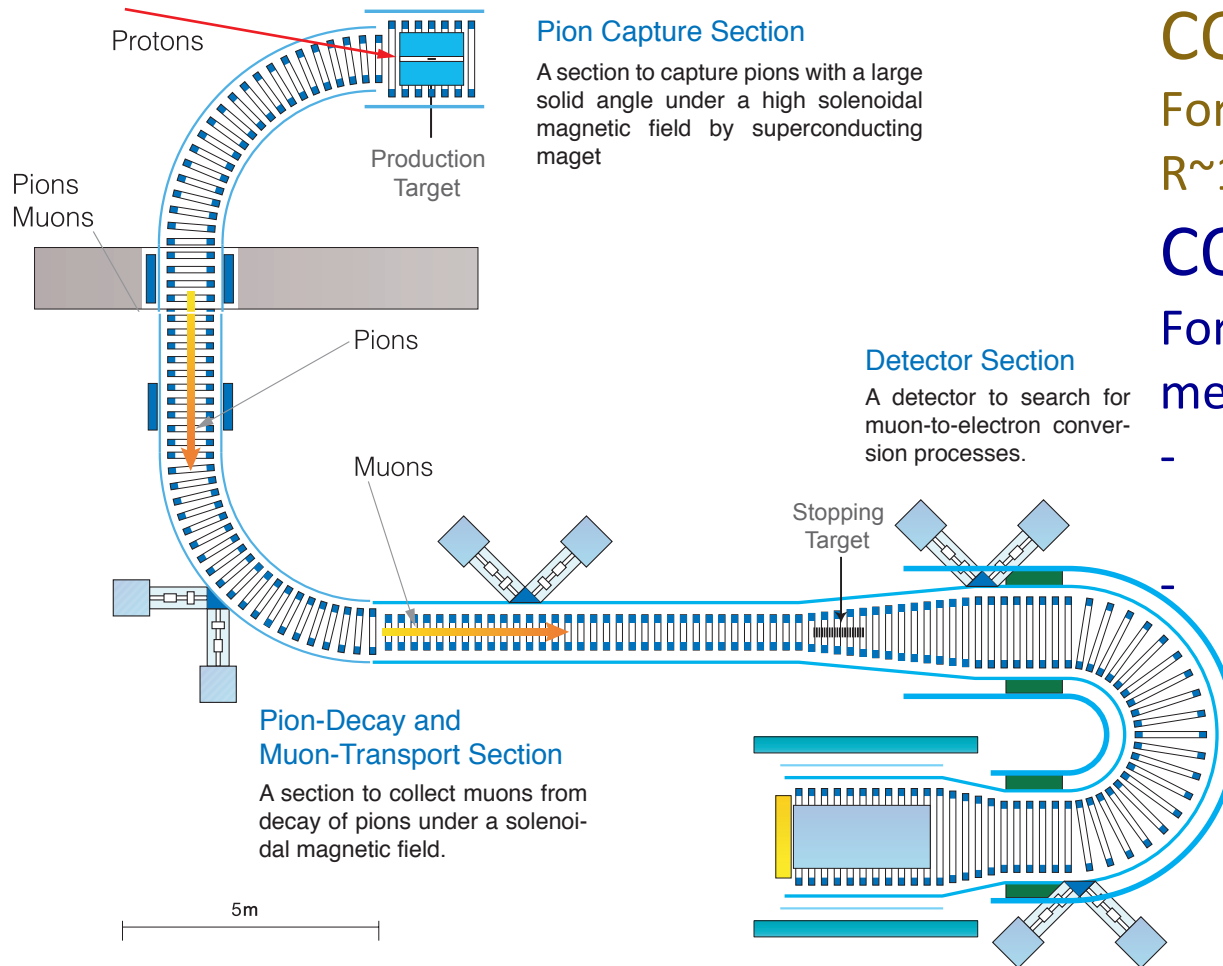
1. Proton hits target and generates pions
2. Negative pions are collected, decays to muon and transported through solenoid
3. Muon stops at target, captured and decays
4. Secondary electrons are detected at detector

COMET Phase-II experiment

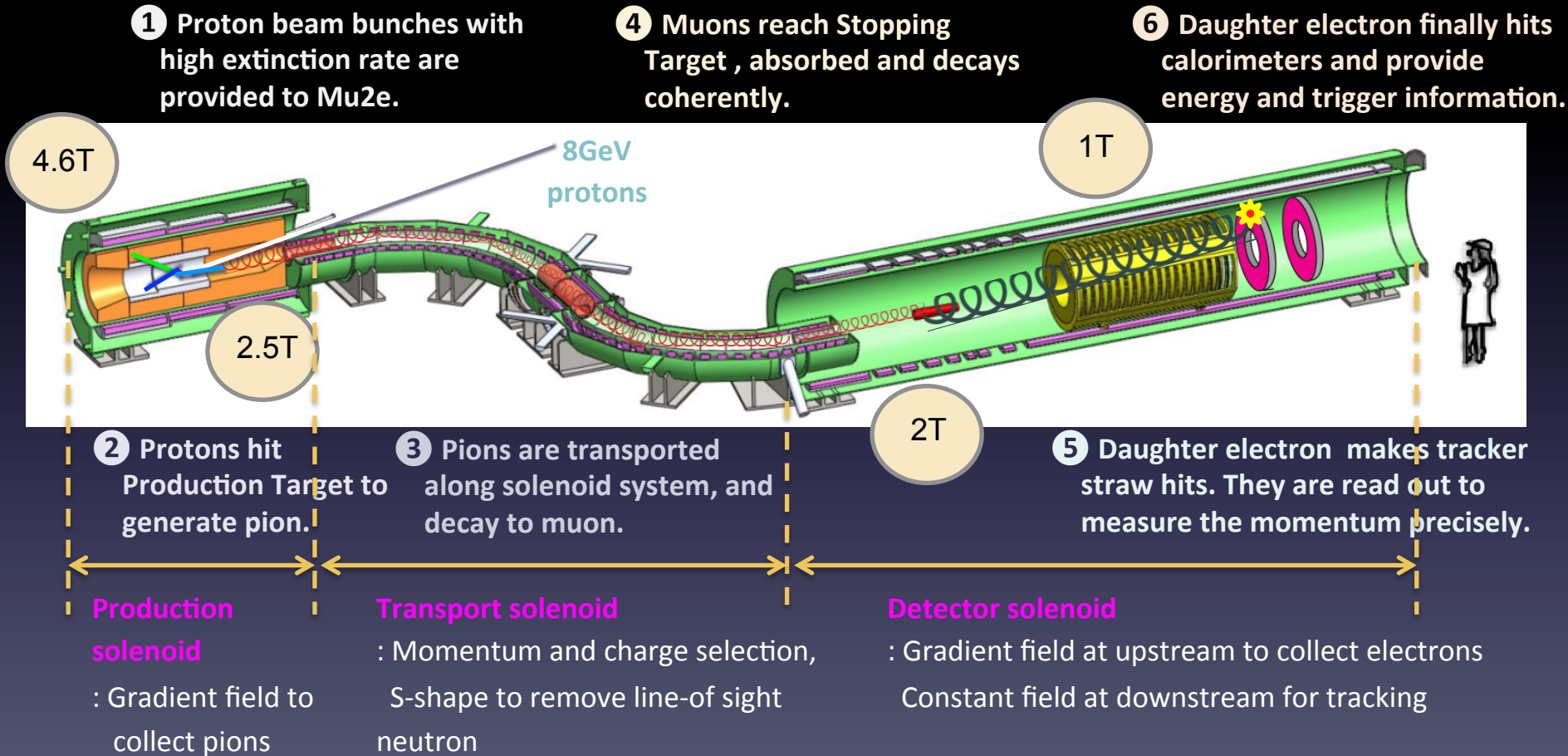
COMET-Phase-I
For BG measurement,
 $R \sim 10^{-15}$ muon conversion

COMET Phase-II
For $R \sim 10^{-17}$ muon conversion
measurement

- proton beam power
3kW $\rightarrow$ 56kW
- x2 DAQ time

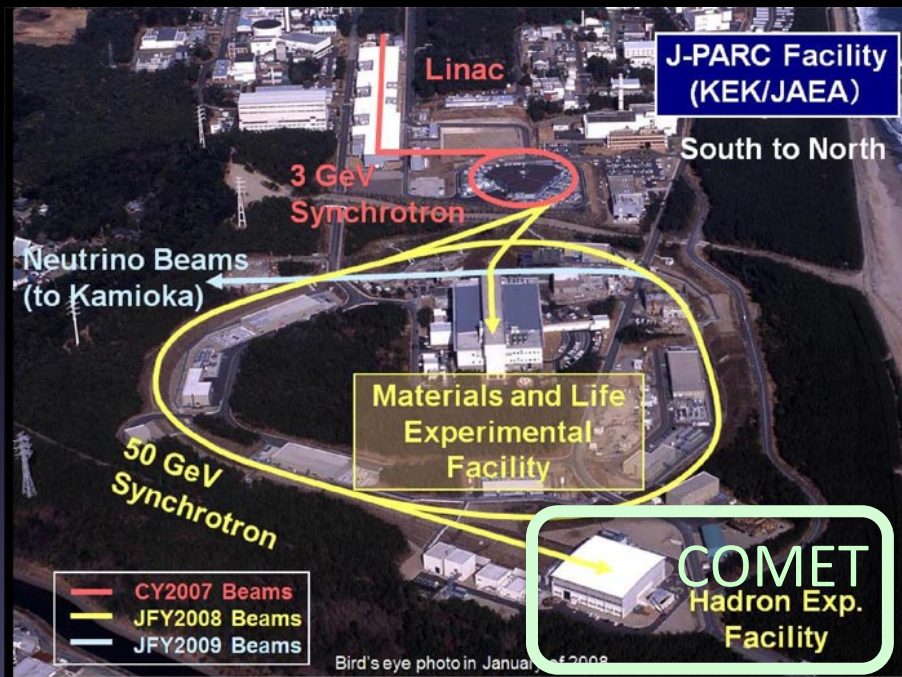


Mu2e experiment

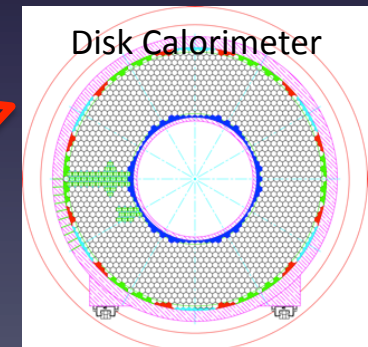
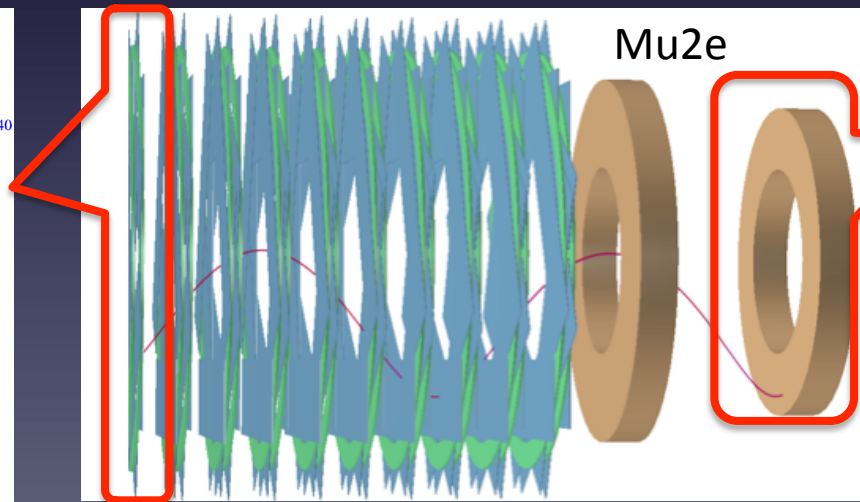
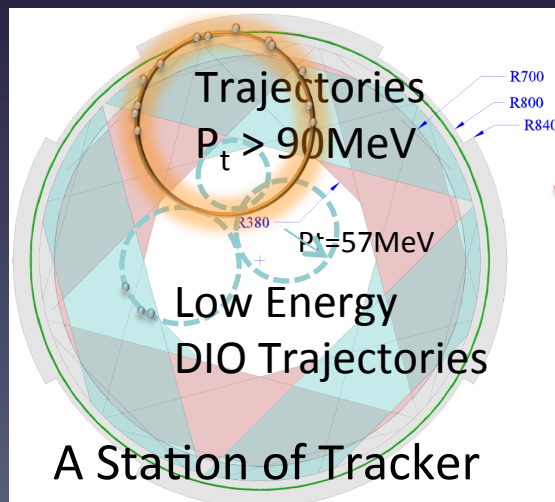
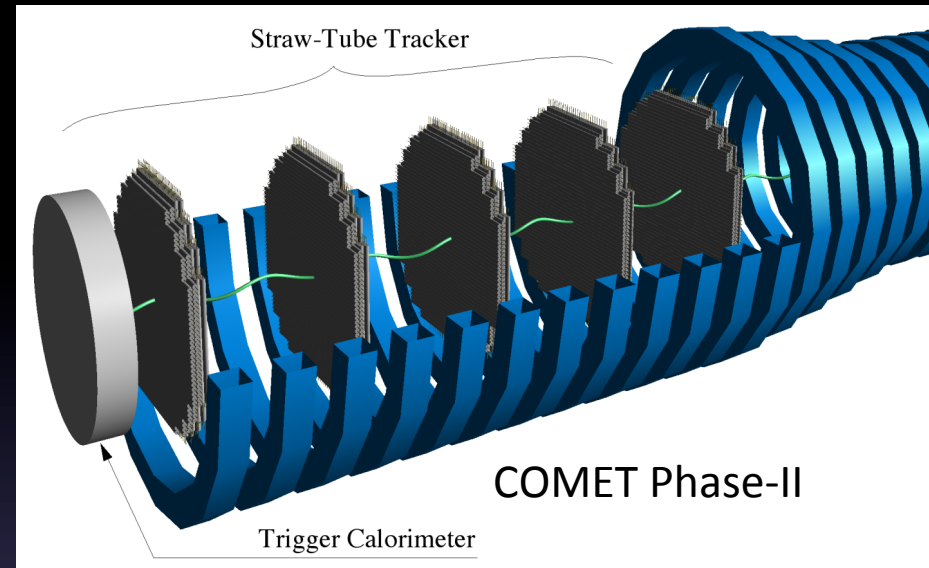
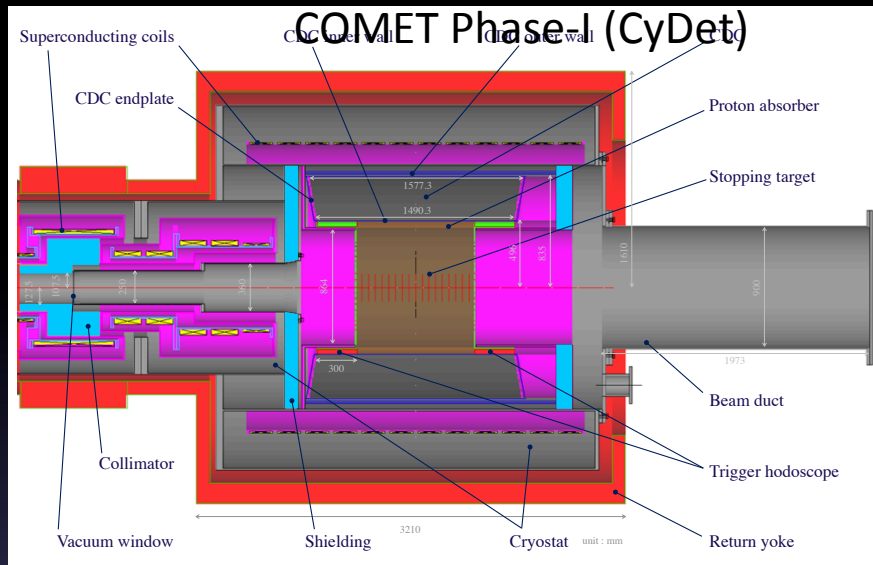


* Not shown : Cosmic ray veto, Muon beam dump, Proton/neutron absorber, Extinction monitor

J-PARC / Fermilab facility



Detectors

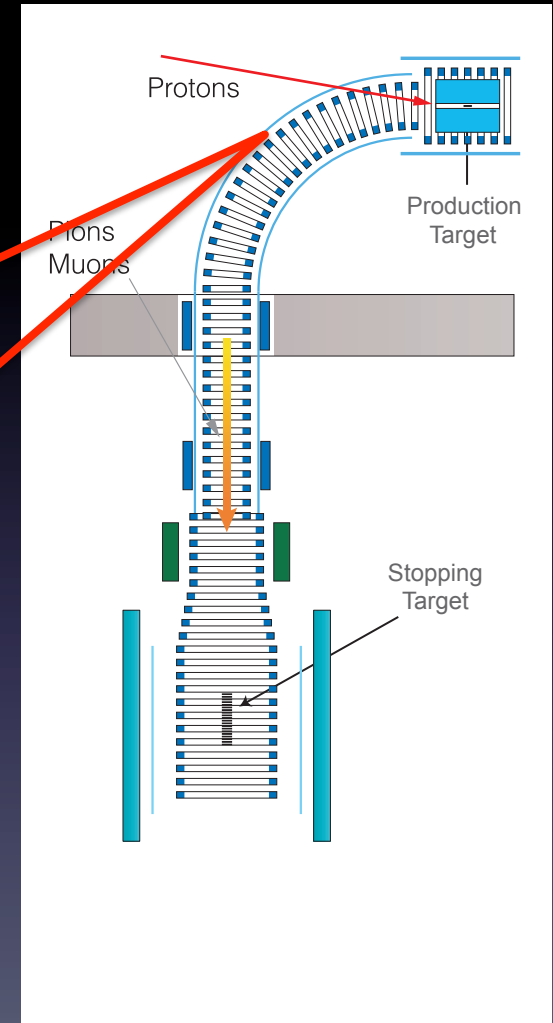
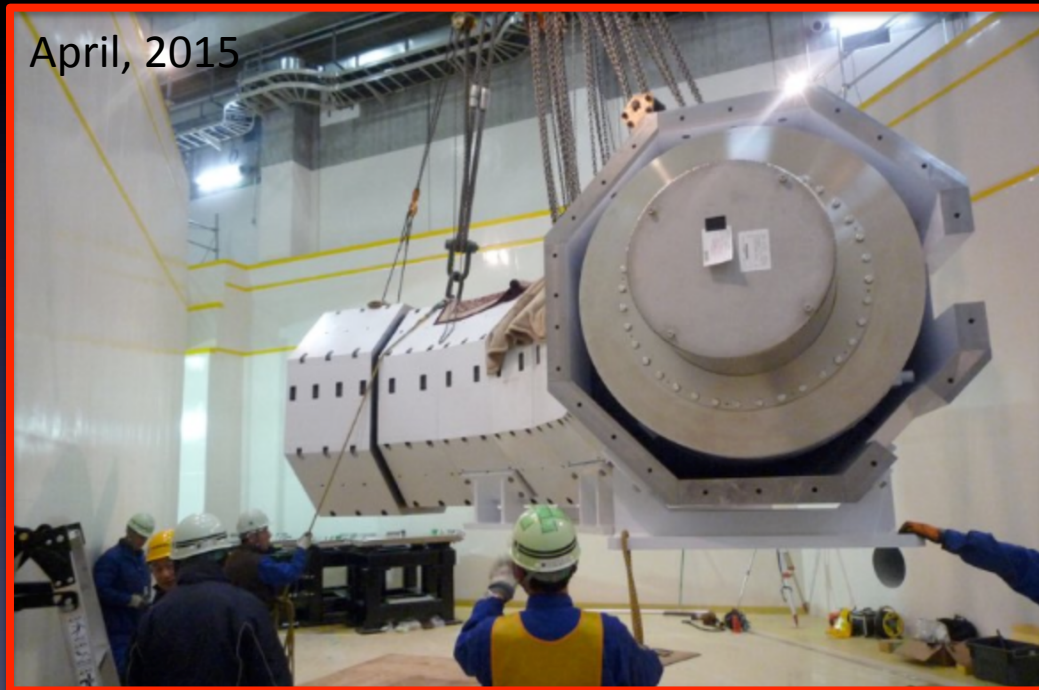


COMET Building Construction

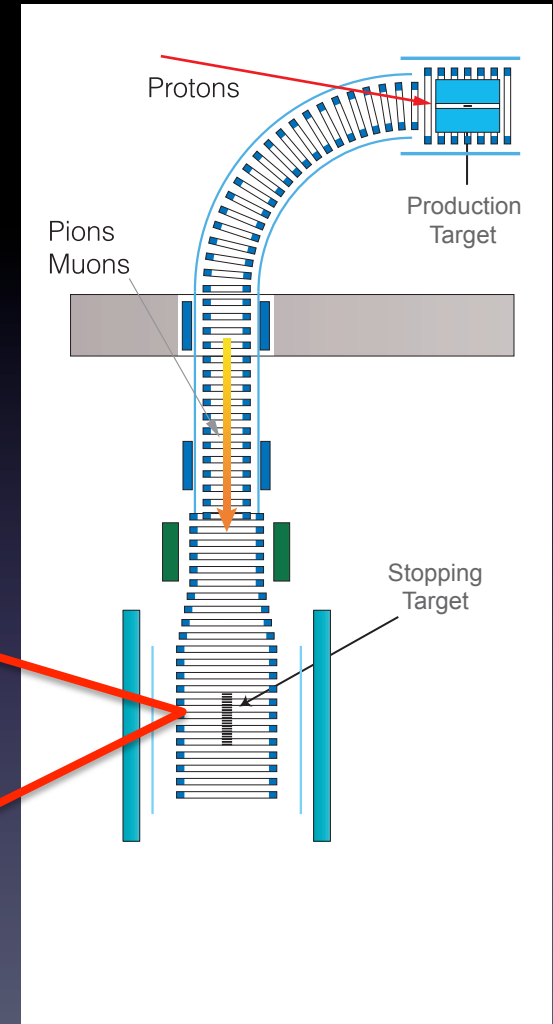
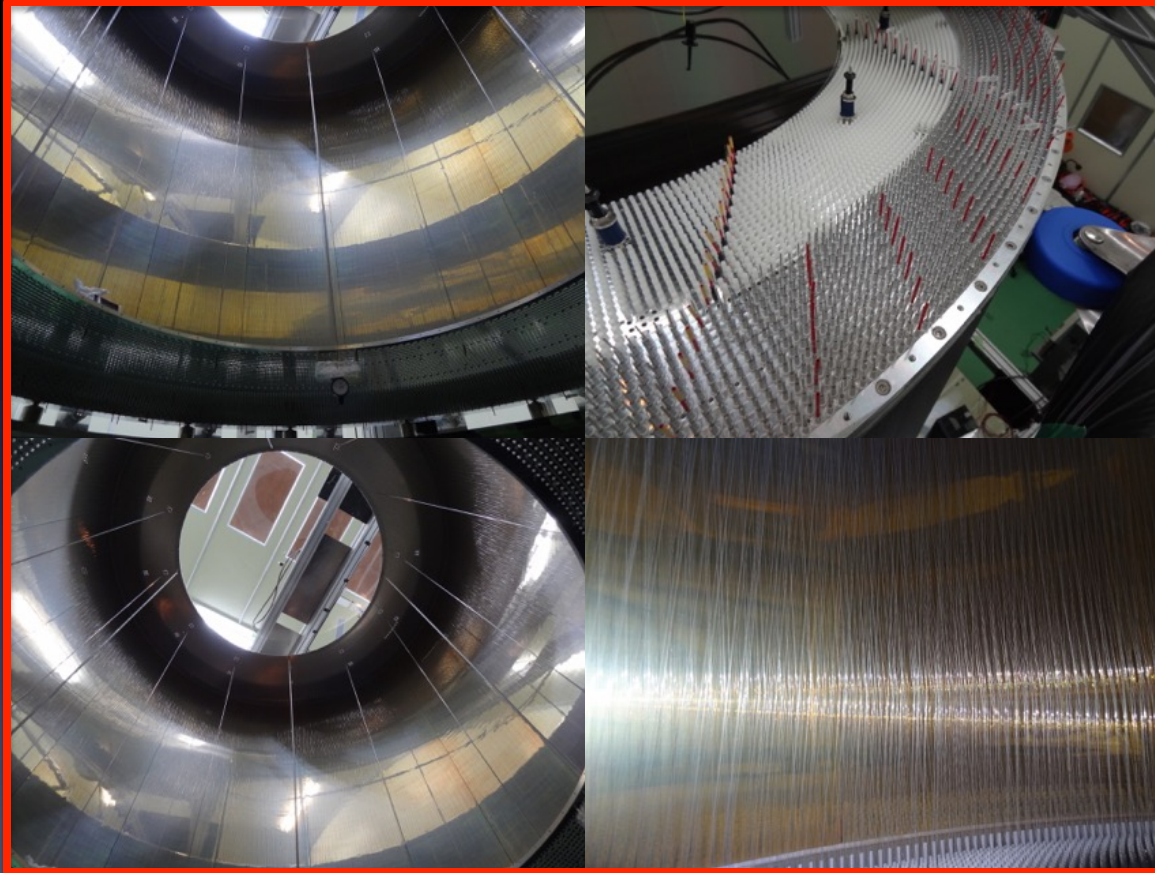
April, 2015



COMET Transport Solenoid



COMET Phase-I CyDet Detector



Mu2e Building Construction

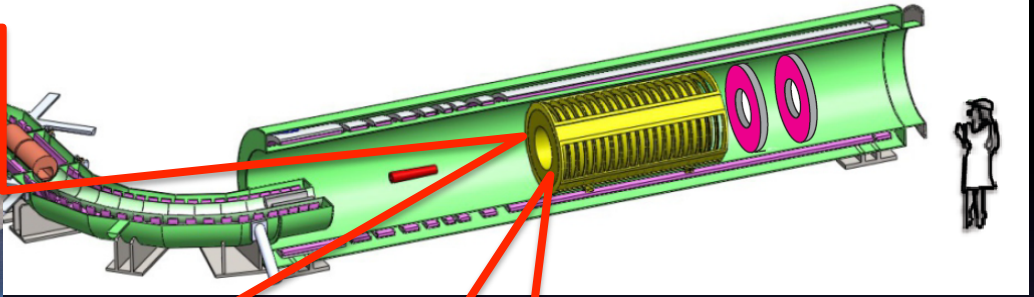


August, 2015
Walls formed

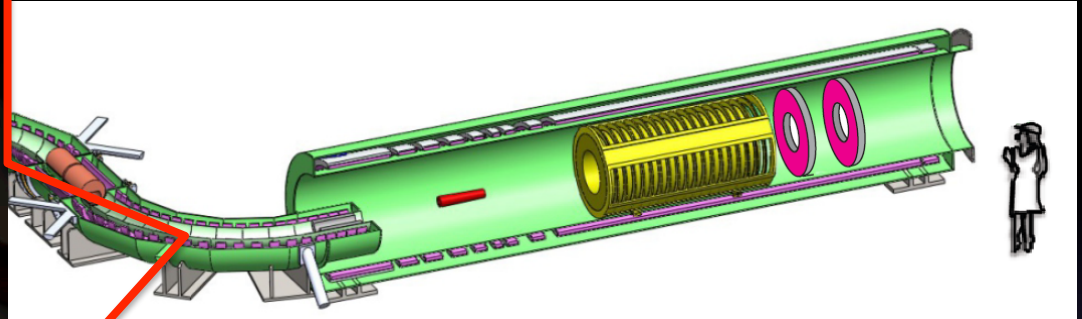
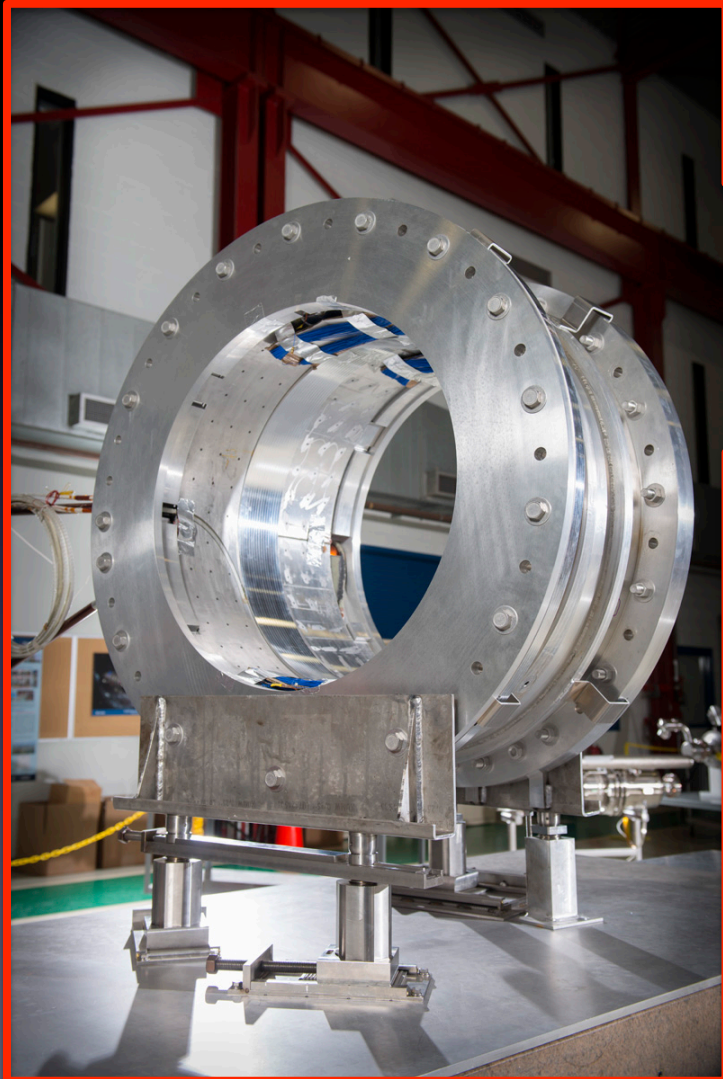


September, 2015
Lower level walls completed

Mu2e Tracker (straw detectors)



Mu2e Transport Solenoid

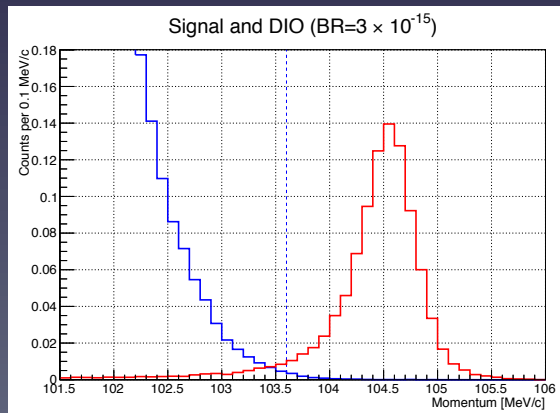


Prototype transport solenoid successfully tested
(Aug 2015)

Acceptance/Background/Sensitivity

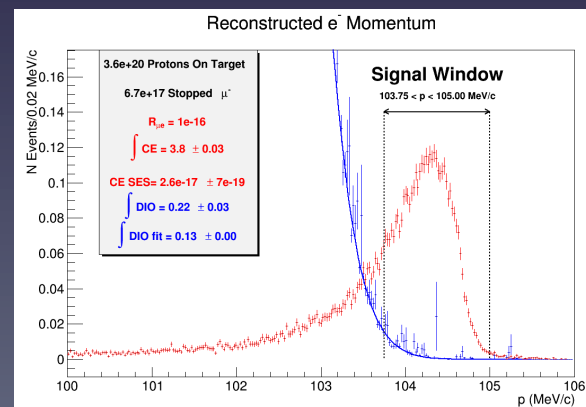
COMET Phase-I

- Net acceptance = 4.3%
 - Geometric acceptance 37%
 - Other detector accep. 64%
 - Reconstruction 15%
- Background = 0.25
 - RPC ~ 0.24
 - DIO ~ 0.01
- SES(Phase-I) $\sim 3.1 \times 10^{-15}$
- SES(Phase-II) $\sim 2.5 \times 10^{-17}$



Mu2e

- Net acceptance = 8.5%
 - Detector acceptance 96%
 - Reconstruction 9.3%
- Background = 0.37
 - DIO ~ 0.20
 - RPC ~ 0.02
 - Anti proton ~ 0.05
- SES $\sim 2.6 \times 10^{-17}$



Schedules

COMET

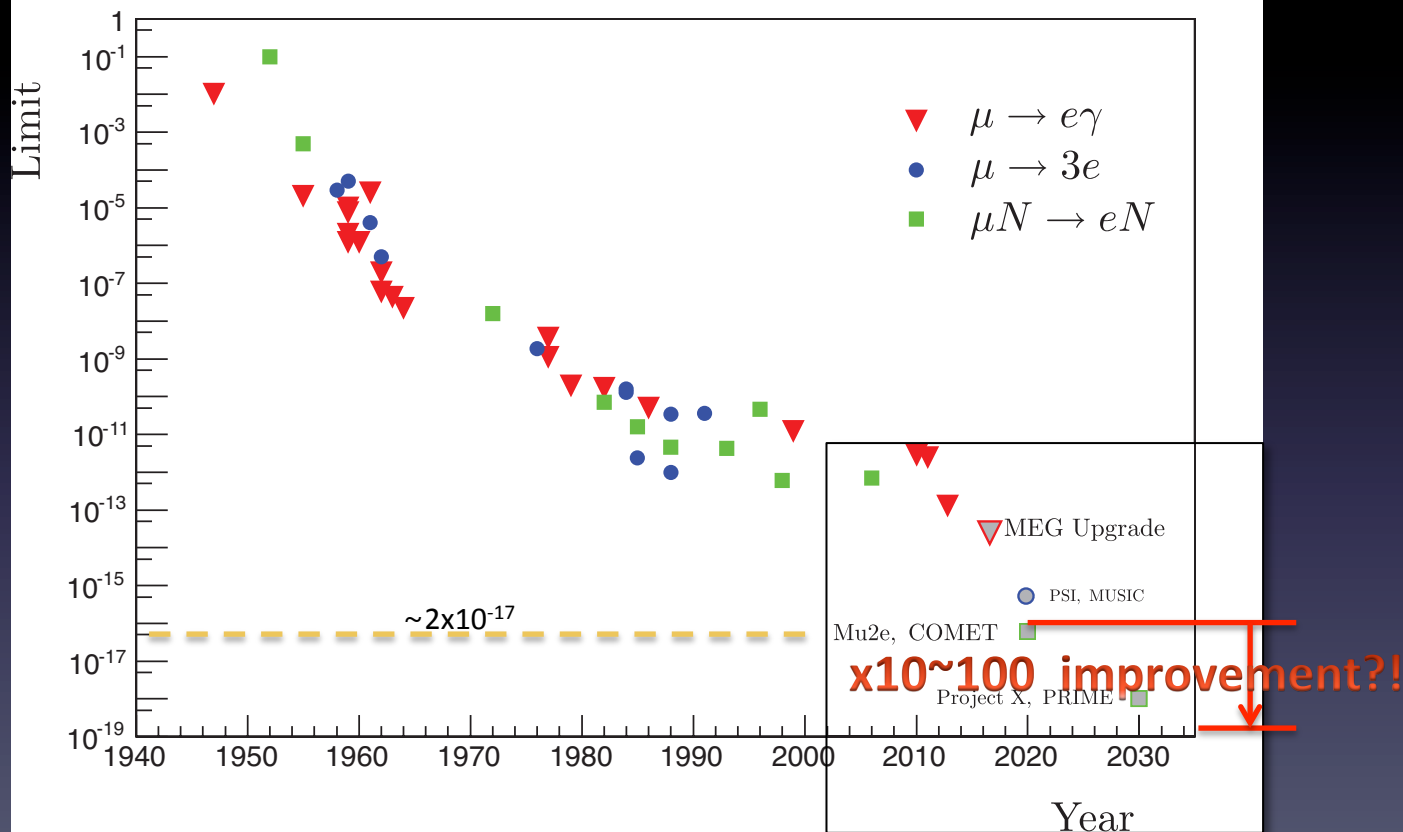
- Proton beam line construction complete in 2017
- Phase-I physics data taking in 2018 or 2019, depending on budget situation
 - Was 2017 or 2018 in original schedule
- Phase-II R&D in parallel with Phase-I R&D and data taking
- Phase-II physics data taking in **2021**
- ~ 1 year data taking

Mu2e

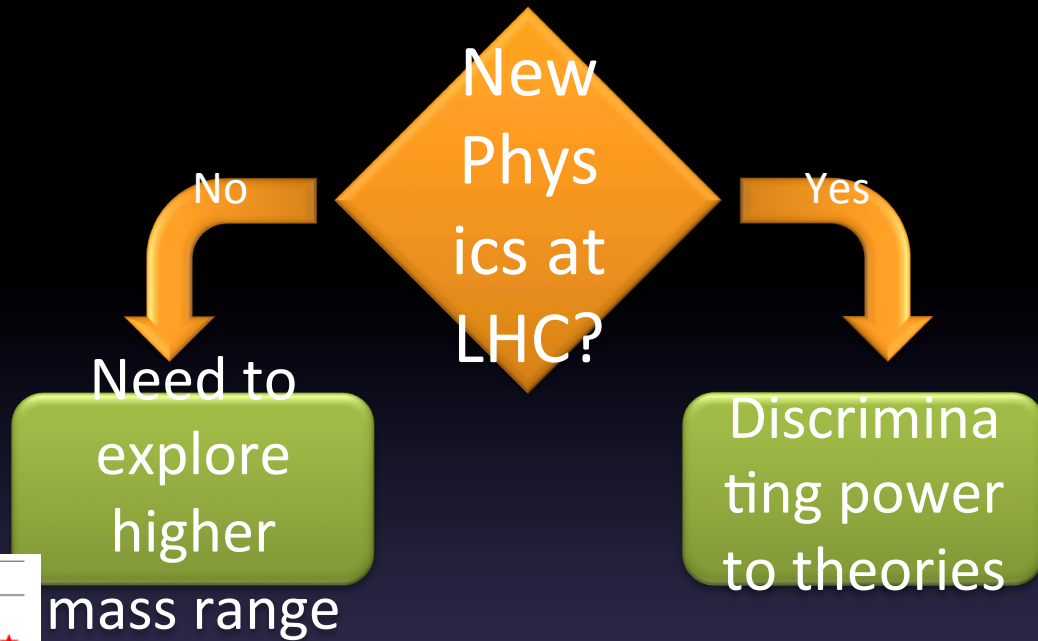
- Strong support by P5 in 2014
- CD (Critical Decision) -3c for approval start of construction in FY2016 2/4
 - CD-3a (long lead procurement) at FY2014 $\frac{3}{4}$
 - CD-2 (baseline) / 3b (phased construction) at FY2015 $\frac{1}{4}$
- Beamline ready at FY2020 4/4
- DAQ start in FY 2021 $\frac{1}{4}$
- ~3 years data taking

Sensitivity

History of $\mu \rightarrow e\gamma$, $\mu N \rightarrow eN$, and $\mu \rightarrow 3e$



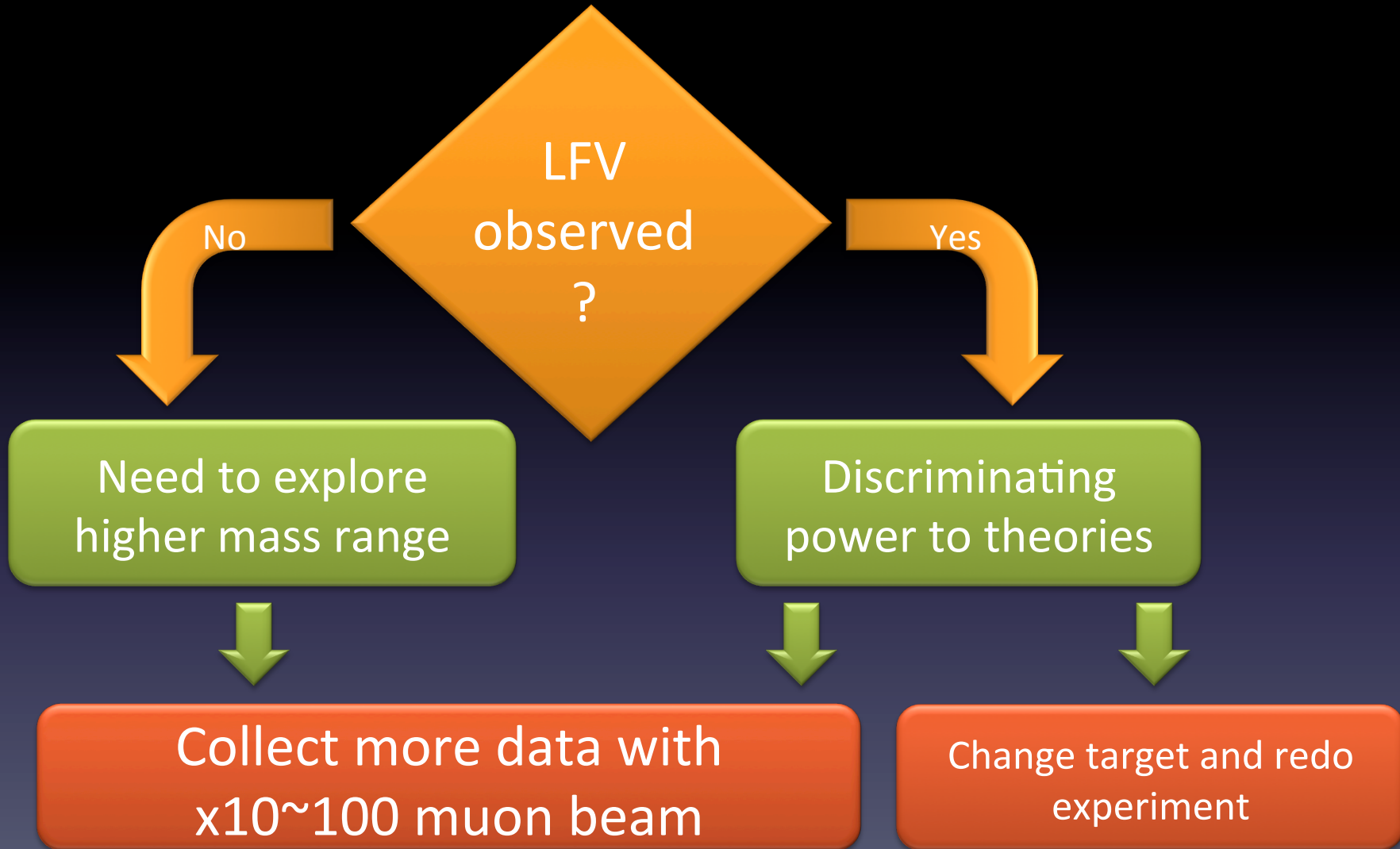
Do we need x100 improvement?



	AC	RVV2	AKM	δ_{LL}	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★	★	★	★	★	★★★	?
ϵ_K	★	★★★	★★★	★	★	★	★★★
$S_{\psi\phi}$	★★★	★★★	★★★	★	★	★★★	★★★
$S_{\phi K_S}$	★★★	★	★	★★★	★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★	★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★	★★★	★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★	★★★	★★★	★★★	★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$\mu \rightarrow e \gamma$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$\tau \rightarrow \mu \gamma$	★★★	★★★	★	★★★	★★★	★★★	★★★
$\mu + N \rightarrow e + N$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
d_n	★★★	★★★	★★★	★	★★★	★	★★★
d_e	★★★	★★★	★	★	★★★	★	★★★
$(g-2)_\mu$	★★★	★★★	★	★★★	★★★	★	?

W. Altmannshofer et al., Nuclear Physics B 830, pp17-94 (May 2010)

Do we need x100 improvement?



Future projects

- Need $> \times 10$ times intense muon beam, $> 10^{12}$ muons/sec
 - Stronger proton beam for more muons
- Keep the current background rate (< 1 BG) even in the increased muon rate
 - Note : $\times 10$ beam does not mean $\times 10$ BG
- **Mu2e-II**
 - Increase proton beam intensity $\times 10$
 - Use Mu2e detector (with minor modification)
- **PRISM/PRIME**
 - Intense muon beam using storage ring muon cooling

Mu2e-II

- Originally studied for Project-X (arXiv:1307.1168)
- Target SES $\sim 2 \times 10^{-18}$ (x10 improvement), BG < 1 event
- Minor modification on detector
 - Such as readout electronics, DAQ
 - To reduce cost
 - However, requires careful redesign on pion production solenoid
- x100 improvement should be “Mu2e-III”
 - Major detector modification
- Can Mu2e-II survive from background?

Mu2e-II @ Project-X

	Mu2e	Mu2e-II @ Project-X	
Proton beam energy	8 GeV	3 GeV	1 GeV
Beam power	8 kW	Up to MW	
Beam width	200 ns	100ns	
Required Proton on Target (POT) (*)	3.6×10^{20}	8.7×10^{21}	4.0×10^{22}
Duty factor	0.32	0.90	
Extinction	$< 10^{-10}$	?	
Muon stops / POT	1.7×10^{-3}	6.7×10^{-4}	1.4×10^{-4}
Muon stops / kW	7.6×10^{16}	8.0×10^{16}	5.2×10^{16}
Required beam power		72 kW	112 kW
Total muon stops	6.1×10^{17} (**)	5.8×10^{18}	
Background (DIO dominated)	0.41	2.34 / 0.79 (***)	

(*) Assuming the same beam time

(**) 9×10^{17} at TDR, depending on simulation setup

(***) Aluminum / Titanium

Ref: arXiv:1307.1168

Mu2e-II @ PIP-II

	Mu2e	Mu2e-II @ PIP-II
Proton beam energy	8 GeV	800 MeV
Beam power	8 kW	8 – 100 kW
Beam width	200 ns	(< 200 ns)
Proton on Target (POT) (*)	3.6×10^{20}	$3.6 \times 10^{21} - 4.5 \times 10^{22}$
Duty factor	0.32	0.90
Extinction	$< 10^{-10}$	$< 10^{-12}$?
Muon stops / POT	1.7×10^{-3}	1.0×10^{-4}
Muon stops / kW	7.6×10^{16}	4.7×10^{16}
Total muon stops	6.1×10^{17} (**)	4.7×10^{18} @100kW

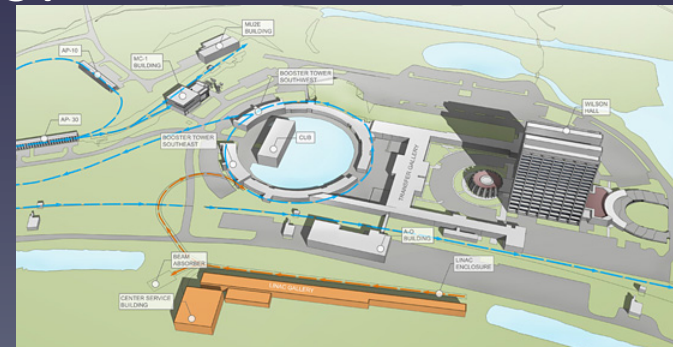
(*) Assuming the same beam time

(**) 9×10^{17} at TDR, depending on simulation setup

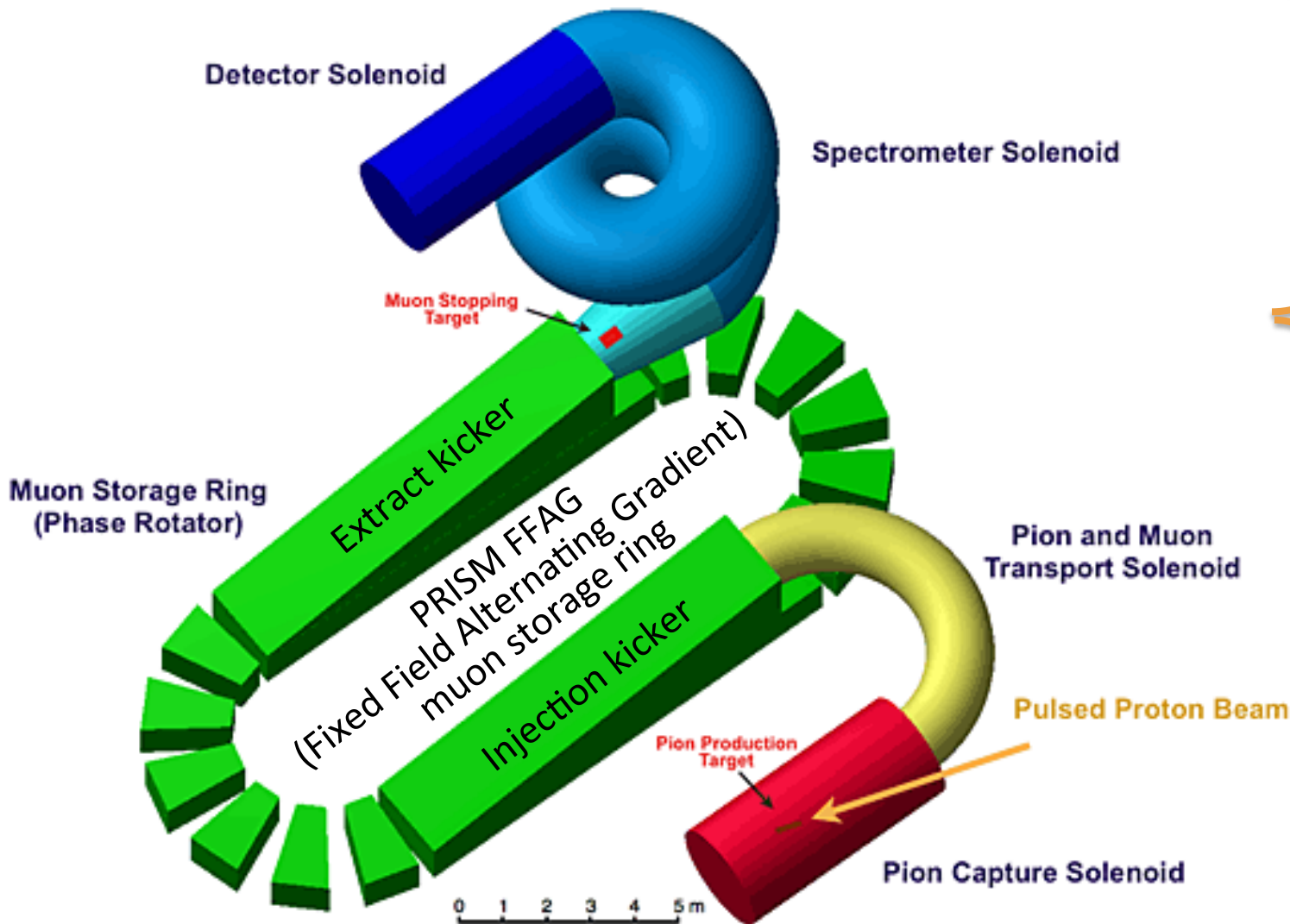
Ref : <https://indico.fnal.gov/getFile.py/access?contribId=1&resId=0&materialId=slides&confId=10235>

Mu2e-II @ PIP-II

- 7.7 times more muons possible for Mu2e-II
 - Not that higher number, detector may survive
- Less beam background because of lower beam energy
 - Ex. no anti-proton background
 - Less radiation effect at pion production
- Background (after reconstruction), detector/electronics occupancy requires careful check
- Driving LINAC beam to muon campus?
 - Not PIP-II original plan



PRISM/PRIME



PRIME

PRISM
Muons to
Electron
conversion

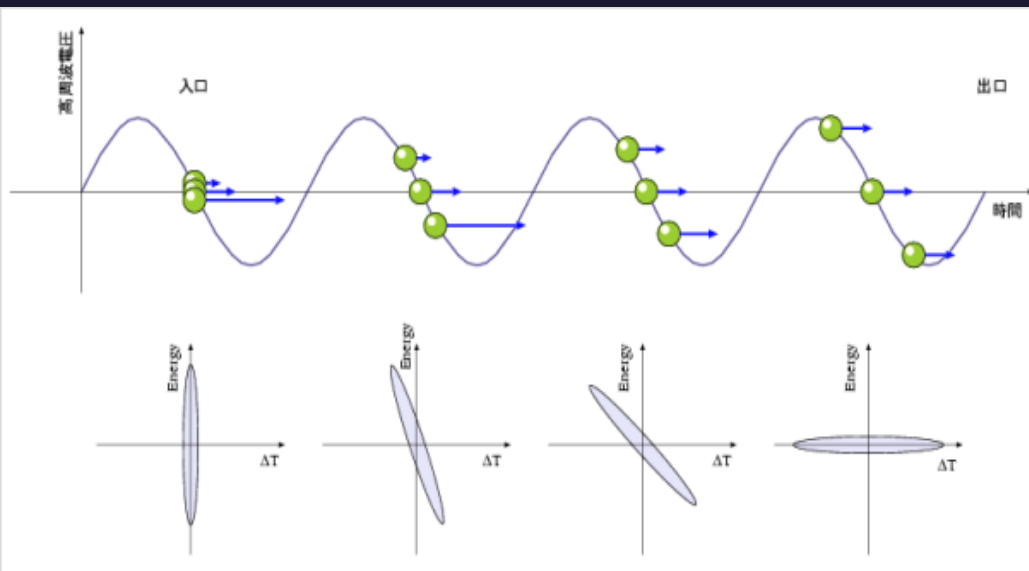
PRISM

Phase
Rotated
Intense
Slow
Muons
source

Ref: <http://prism.phys.sci.osaka-u.ac.jp/research/r003.html>

FFAG Muon Storage Ring

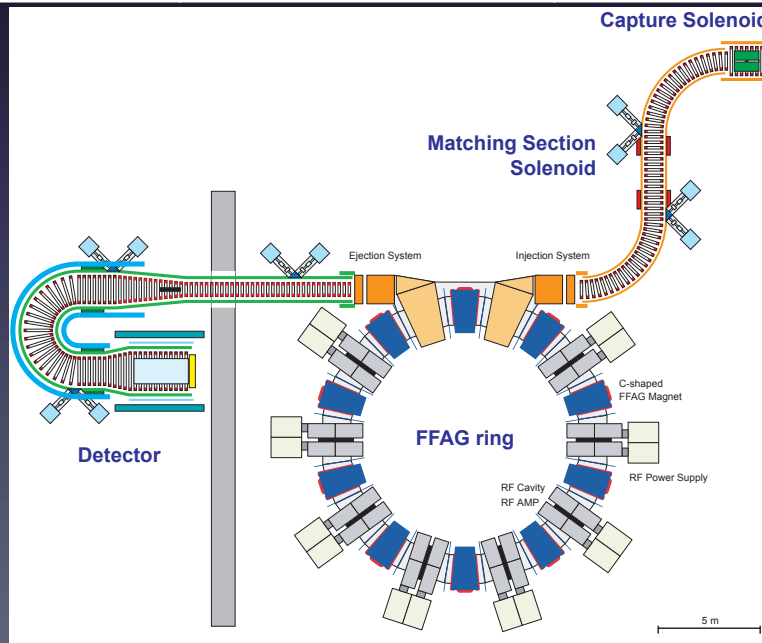
- To get more **intense, monochromatic, pure** (from pion) muon beam
- Use Alternating electric field
 - Faster muon lose energy, slow muon get energy
- Storage ring
 - Pion decays away



Ref: <http://prism.phys.sci.osaka-u.ac.jp/research/r003.html>

COMET vs PRISM/PRIME

Specification	COMET	PRISM/PRIME	Comments on PRISM/PRIME
Muon intensity	$\sim 10^{11}$ muons/sec	$10^{11} - 10^{12}$ muons/sec	$\sim 10^{11}$ muons/sec
Muon momentum	0-140MeV (70MeV peak)	68 ± 2 MeV/c	3% spread (Enhanced from 30% by phase rotation) Less than 70MeV to suppress DIF
Beam repetition	~ 1 MHz	100-1000Hz	Due to kicker Better to reject cosmic ray origin BG



Ref: Y.Kuno, Mu2e workshop (2016)

How to get x100 improvement?

Ref: Private communication, Y.Kuno @ Osaka

- $\times(1/2)$ from reduced beam acceptance from solenoid to FFAG
- $\times 3$ from removing detection time window (no pion)
- $\times 3$ from pion capture improvement
- $\times 20$ from 56 kW $\rightarrow$ 1MW

Multi MW proton source is critical!

Instead of Summary: Much less chance than “needle in haystack”

Courtesy to B.Bernstein @ FNAL



- Or, handful of grains out of earth beach
 - Earth coastline length = $\sim 800,000\text{km}$
 - Assuming width = 50m, depth = 1m, grain/cm³ ~ 8000
 - Total grains in earth beach $\sim 3 \times 10^{20}$
- $\text{SES} \sim 10^{-19}$ is a few finger tips out of earth beach
- MW proton source is critical to explore $O(10^4)$ TeV energy scale